Seasonal and Regional Biases in Aquarius V4:
Geophysical Model

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Background

- Last telecons:
  - V4.1 ("instrument only calibration") shows spurious annual signal (+ drift after October 2014 in some of the channels).
    - Most likely instrument related.
    - Visible in cold sky calibration in most of the channels.
- Go back to V4.0 (ocean calibration, global HYCOM)
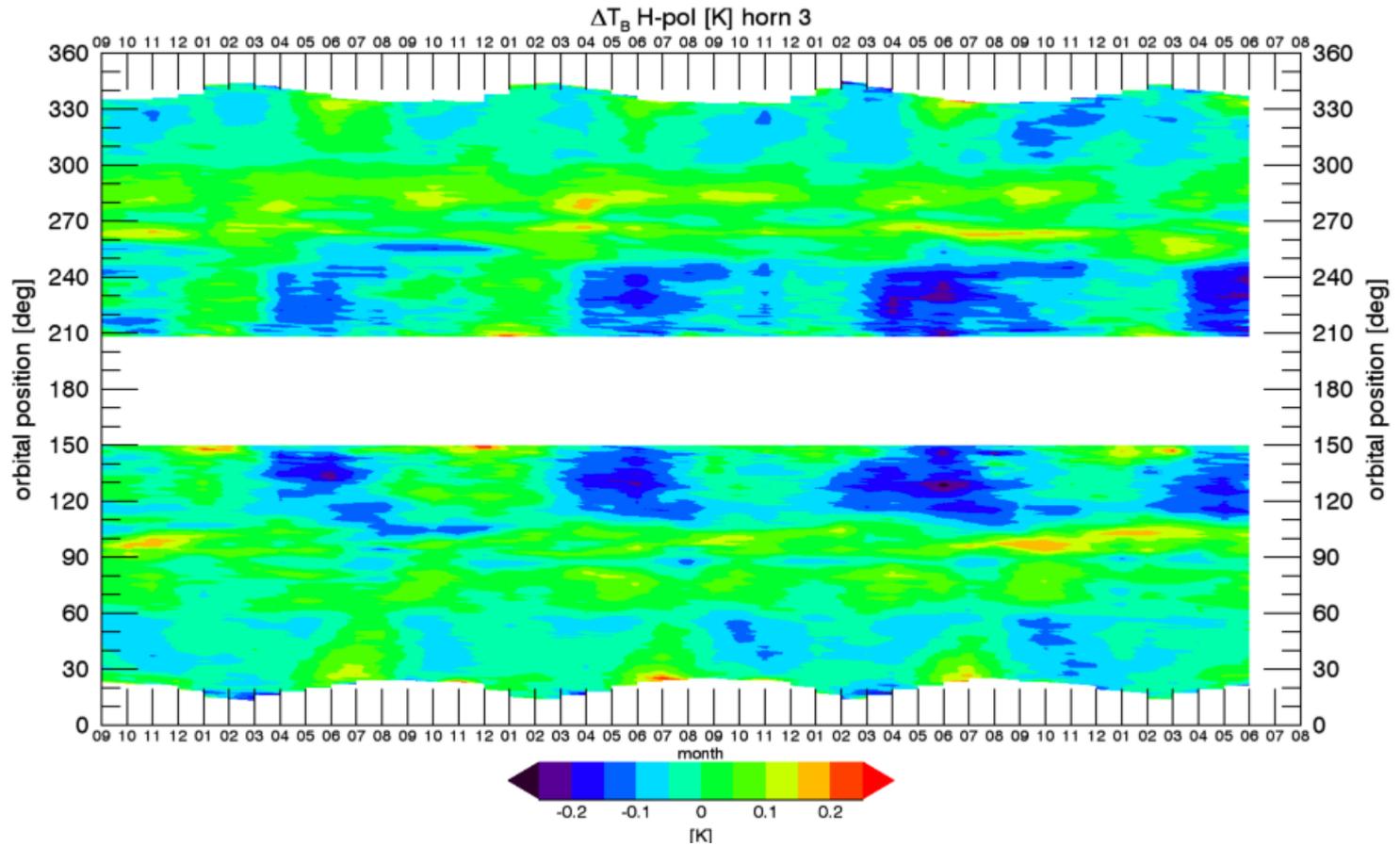 - Analyze seasonal and regional biases.

Salinity
V4.0

ΔSSS [psu] horn 3
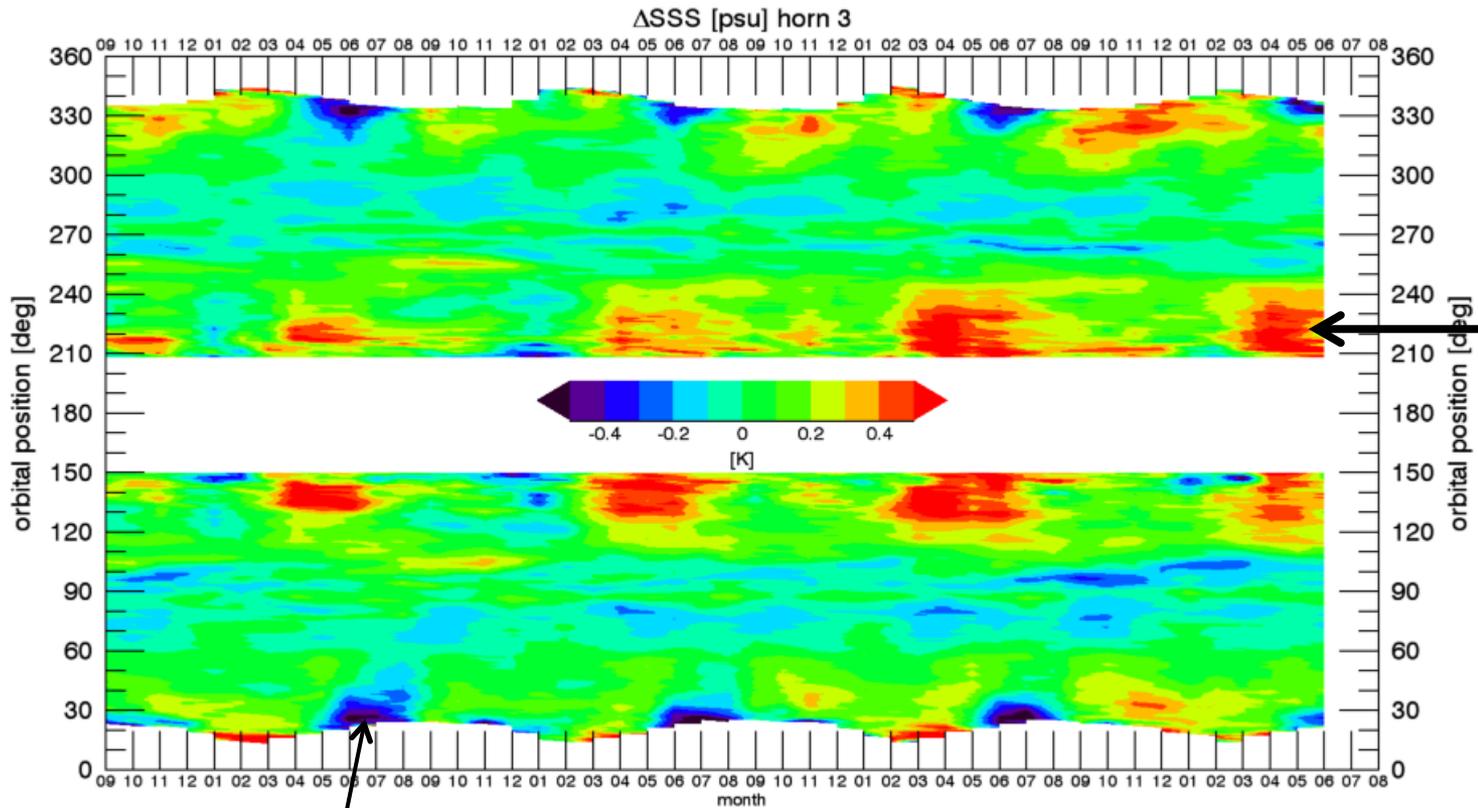
galaxy? high winds?
strong in both asc/dsc
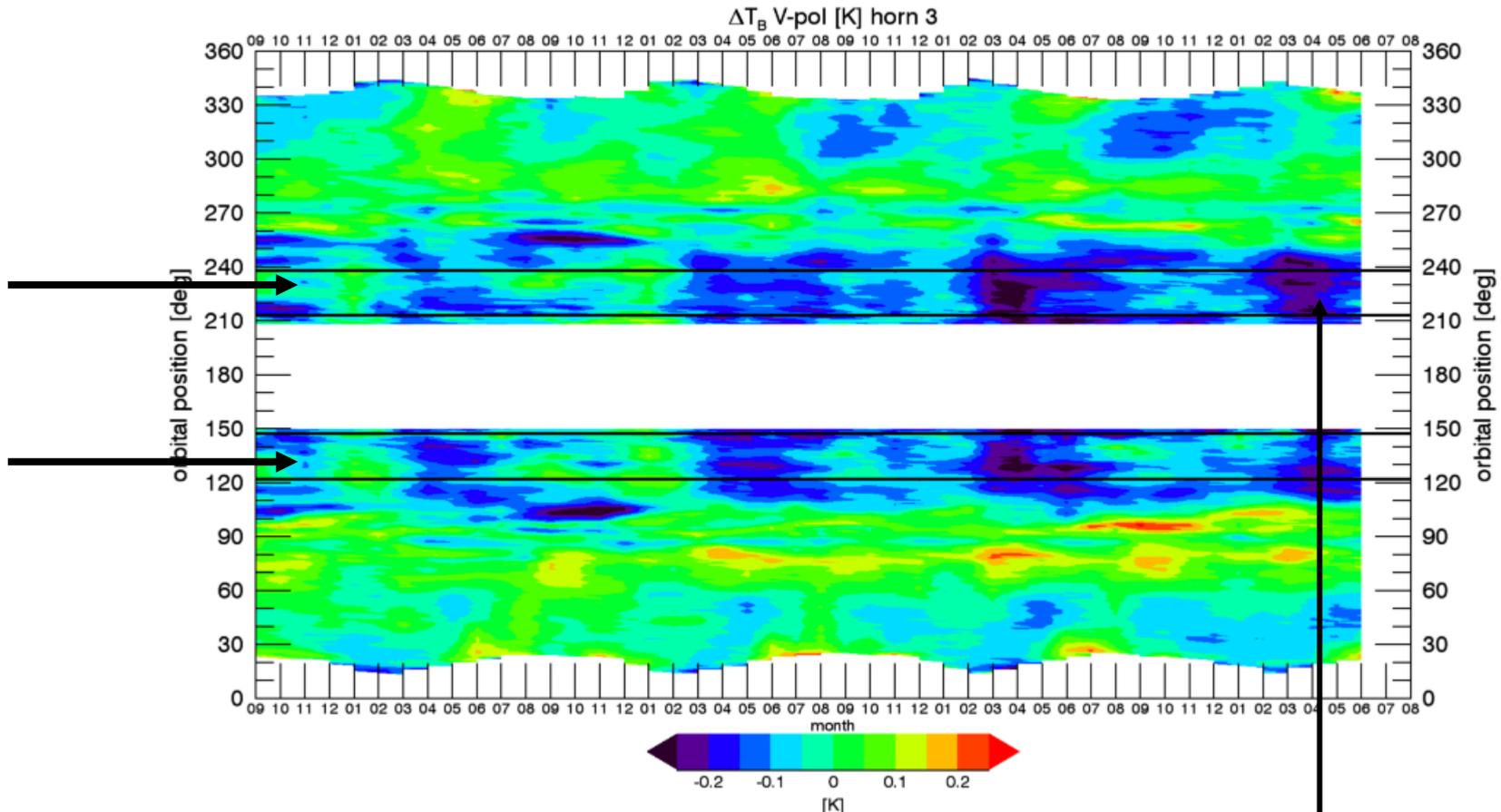
?

TB v-pol
measured - calculated
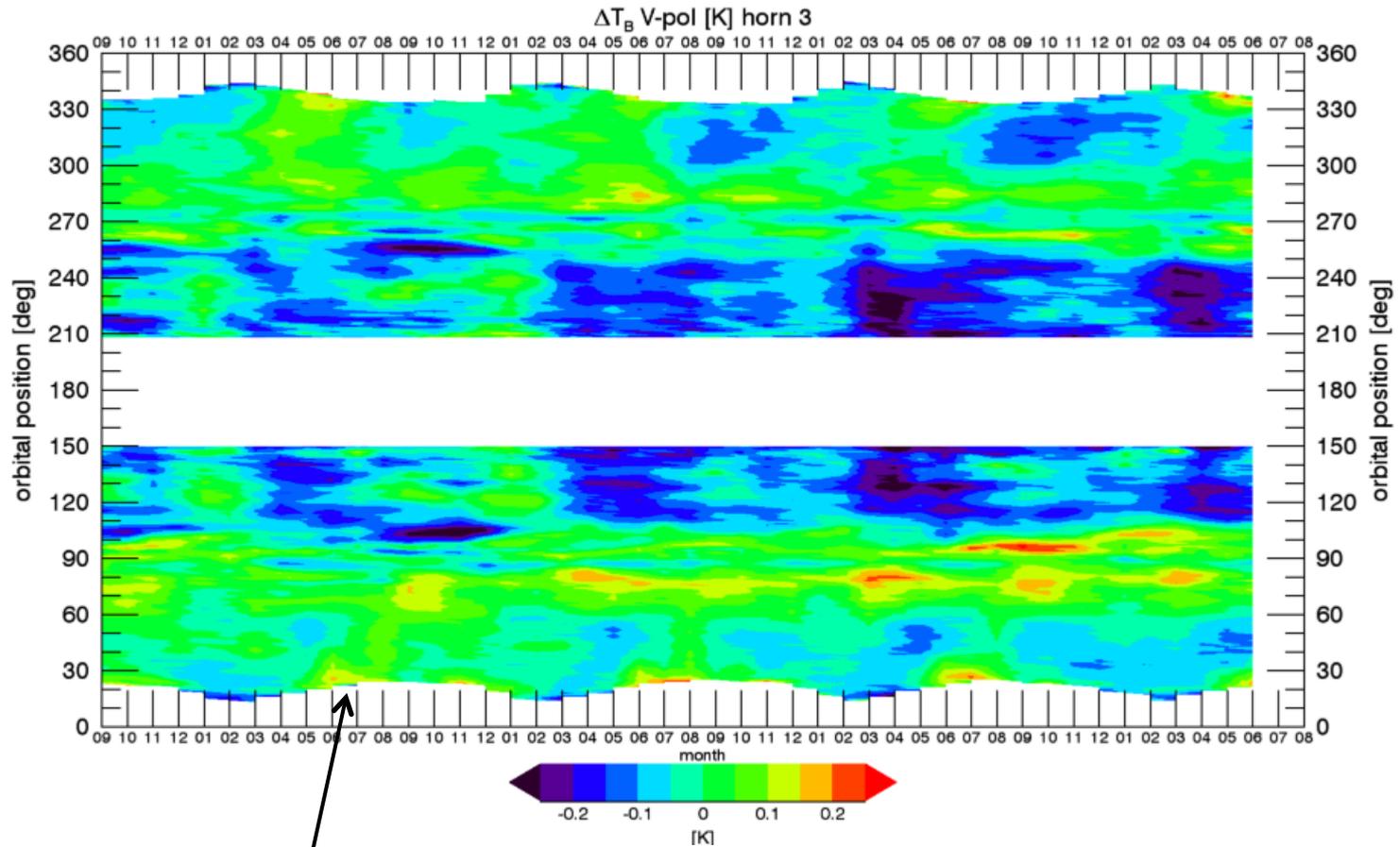
galaxy?
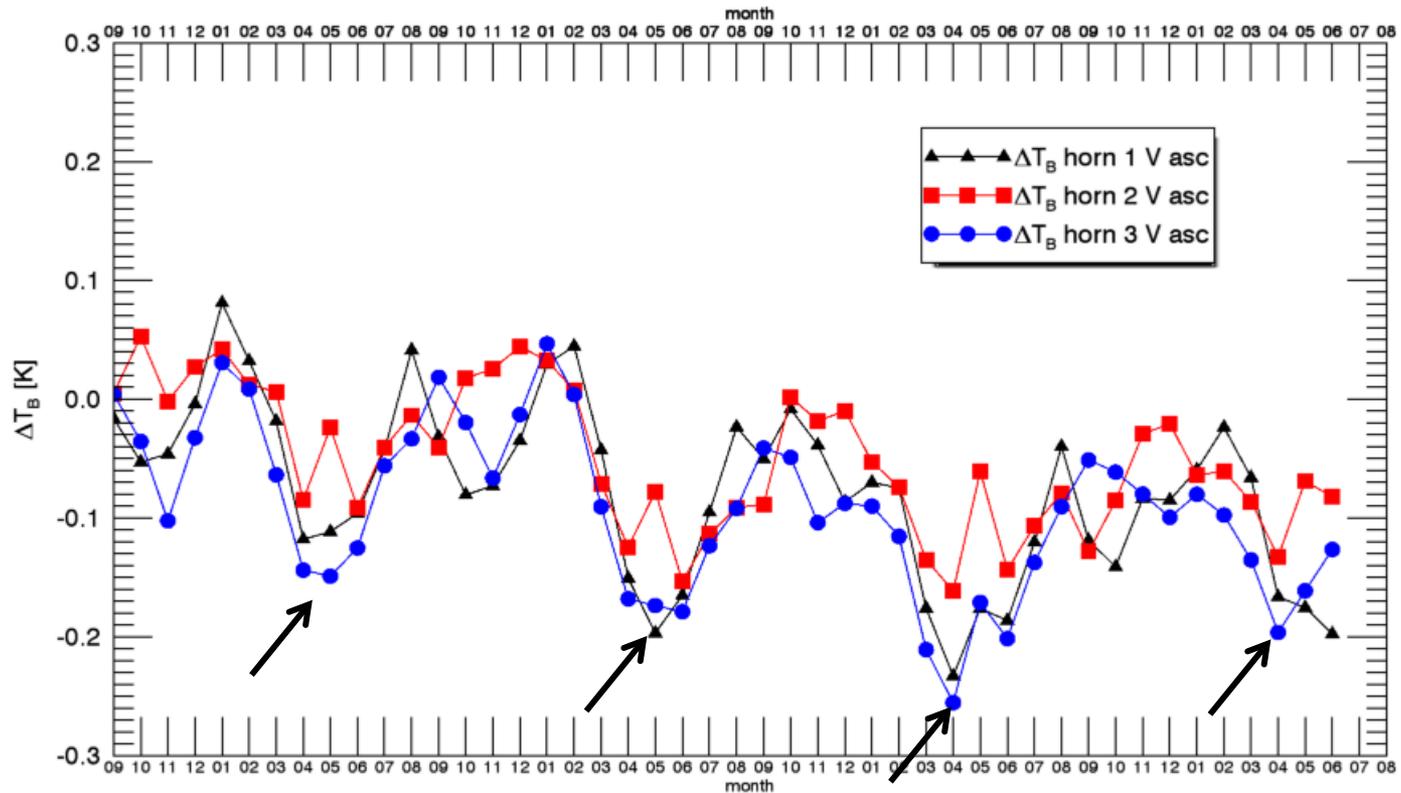Radiometrically small
cold water !

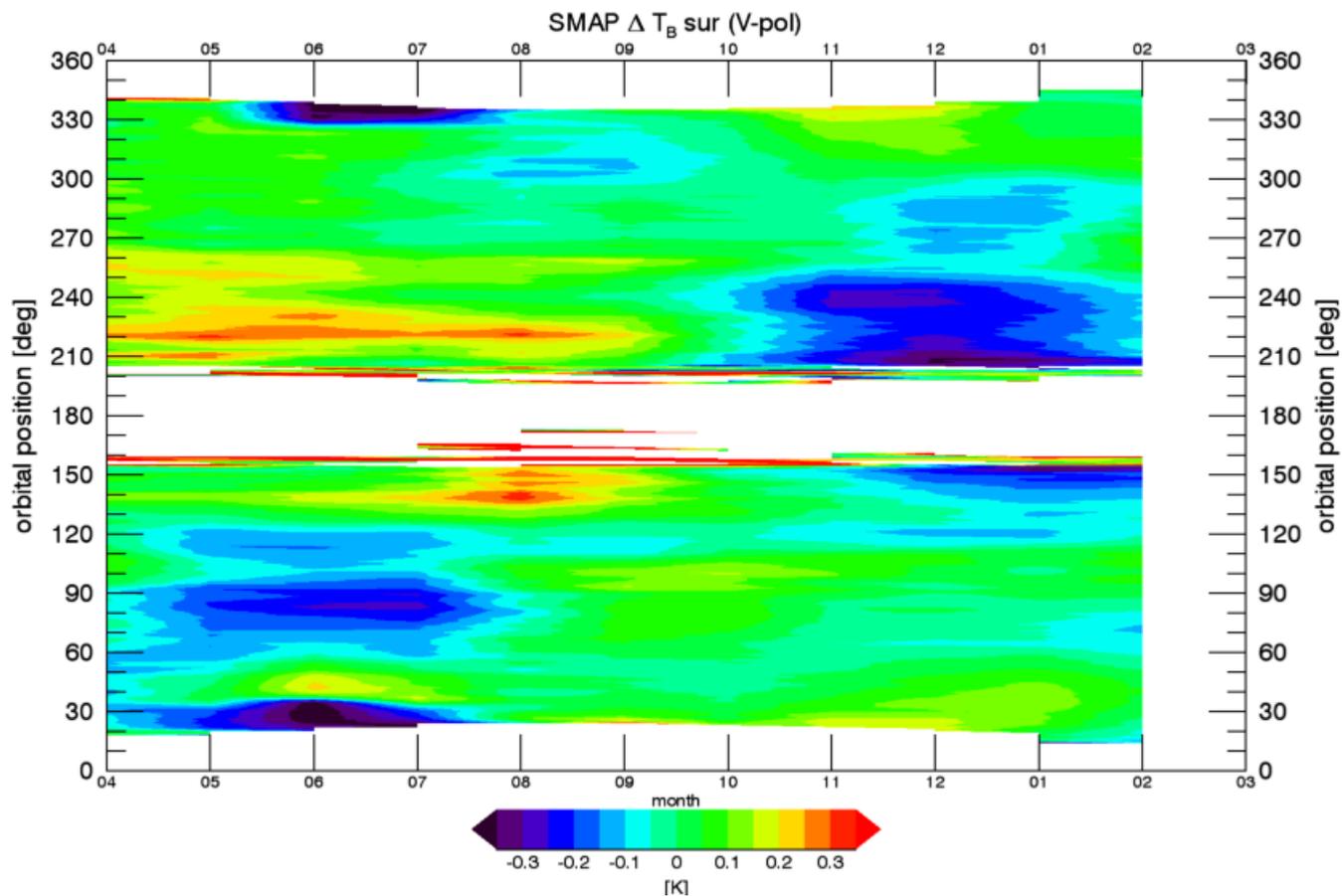# TB h-pol

Time Series: Peak zonal average

V-pol > H-pol
only small asc/dsc differences
intra-annual difference
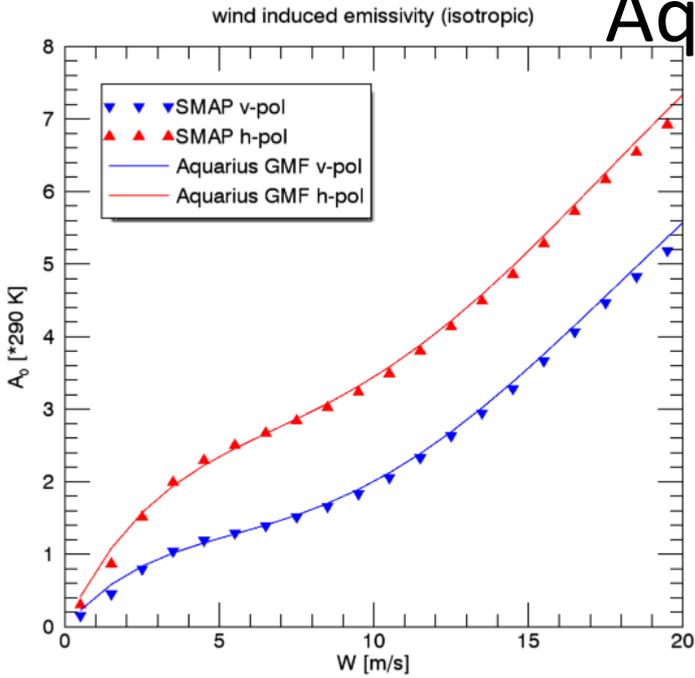
Inter-Beam Differences

horn1, horn3 > horn 2
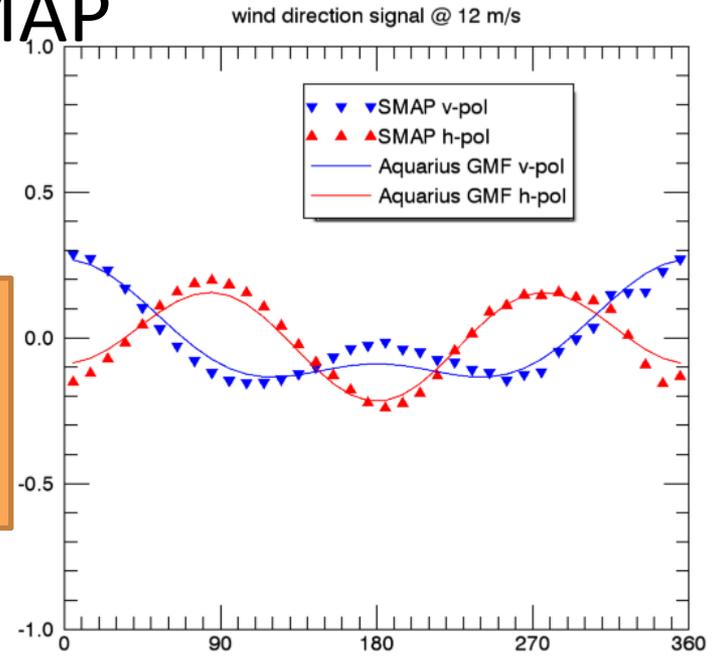peak in MAM

SMAP

SMAP Δ T_B sur (V-pol)

Seasonal and regional biases in SMAO are largely caused by emissive reflector and the deficiencies of the thermal model.
There is no resemblance in SMAP to the Aquarius biases in the N hemisphere during MAM.
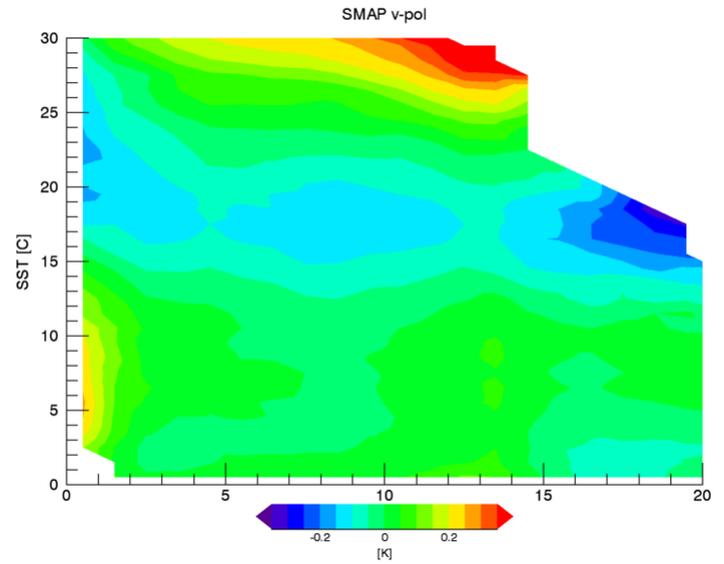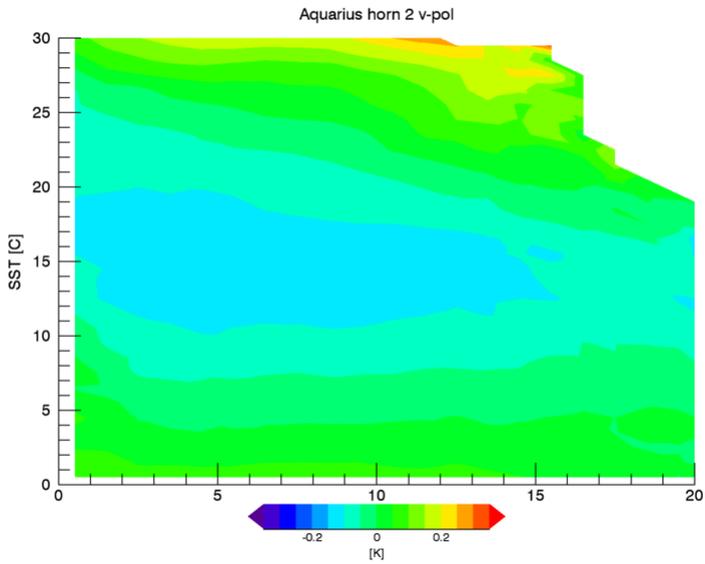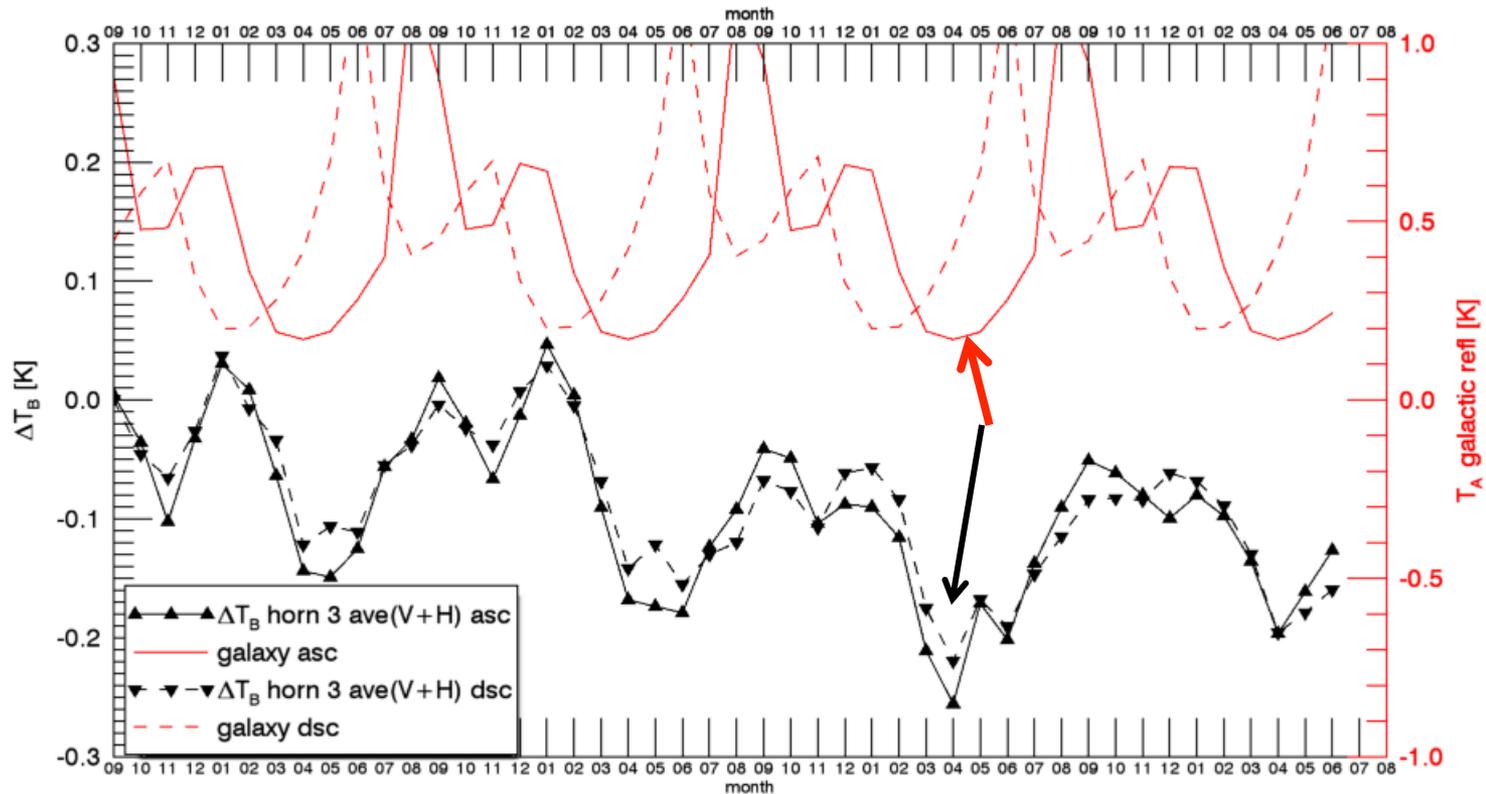
Aquarius vs SMAP

good consistency between Aquarius and SMAP roughness model

SST bias adjustment

Galaxy

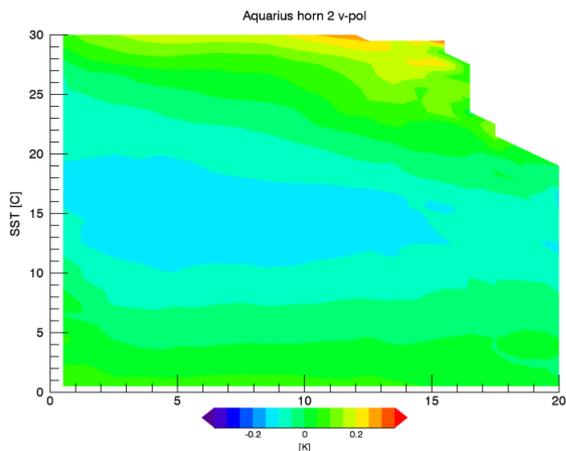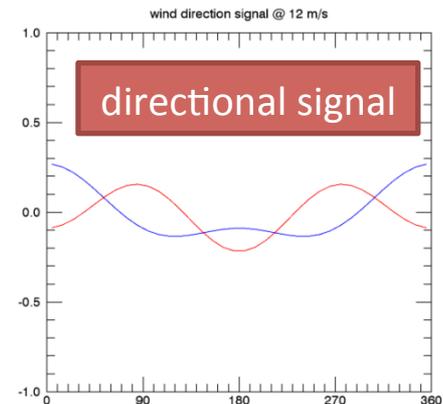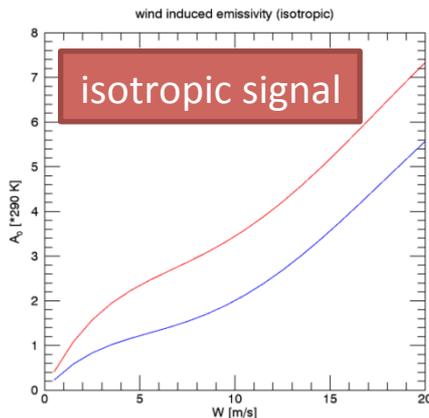strong ascending/descending signature in galaxy
not observed in TB
very small values for TA gal refl at the times when we observe the large TB biases

Surface Roughness Correction
In a nutshell
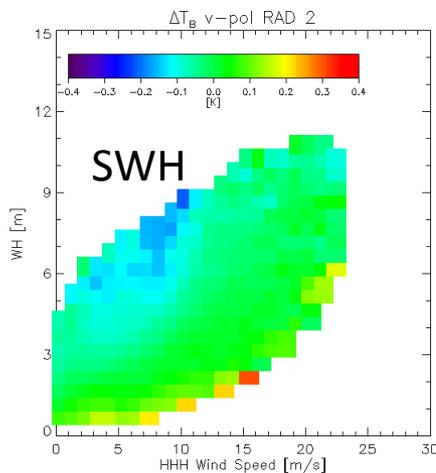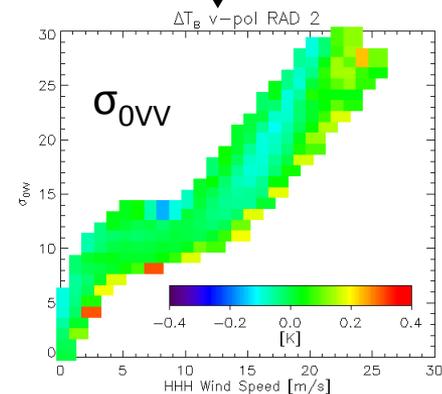
Residual SST/WSPD Dependence
global

residual W/SST crosstalk v-pol horn 3
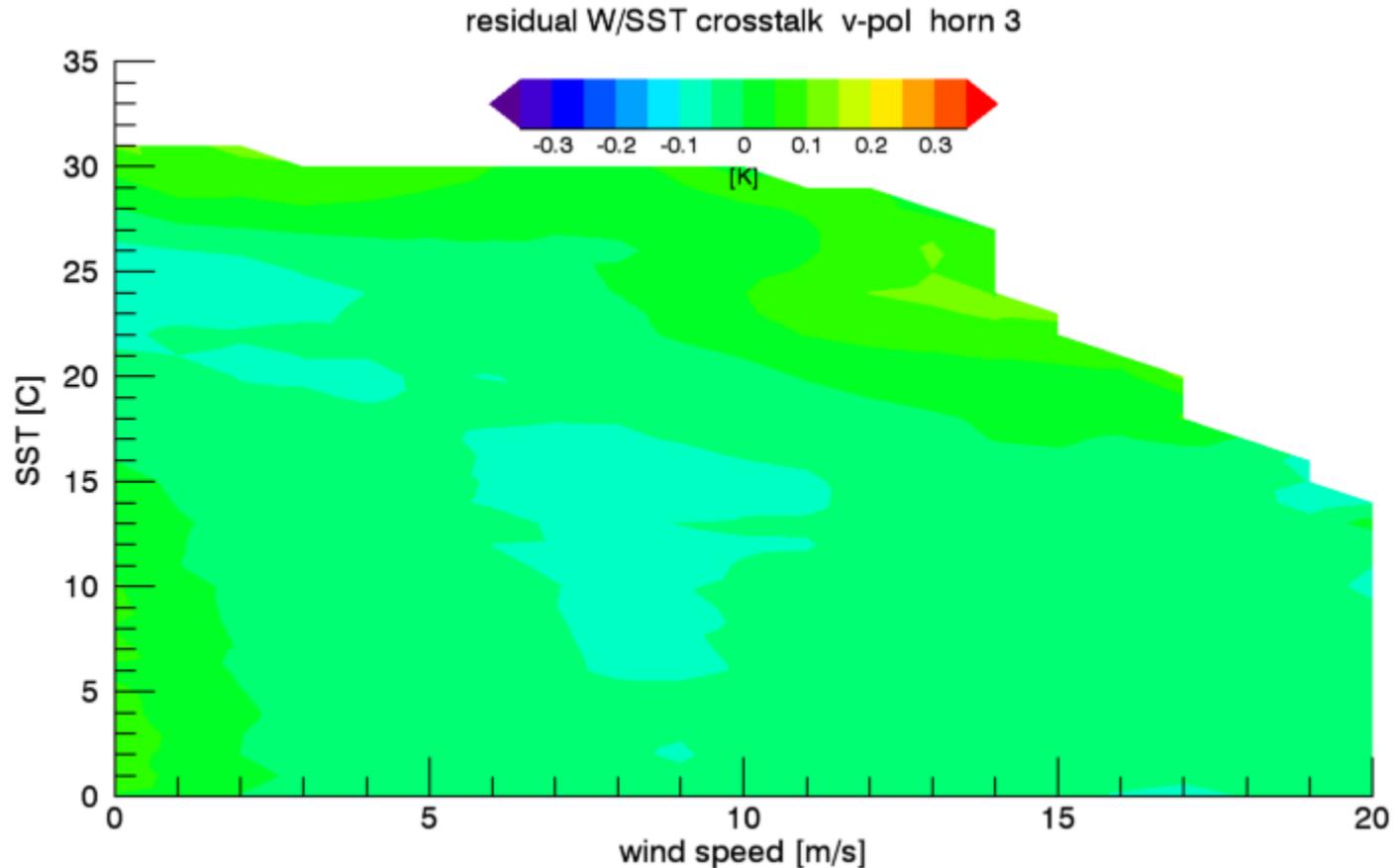
very small – none.
The SST bias adjustment in the roughness correction does what it is supposed to do.

Residual Wind Direction Dependence
global

residual wind direction crosstalk @ 10 m/s
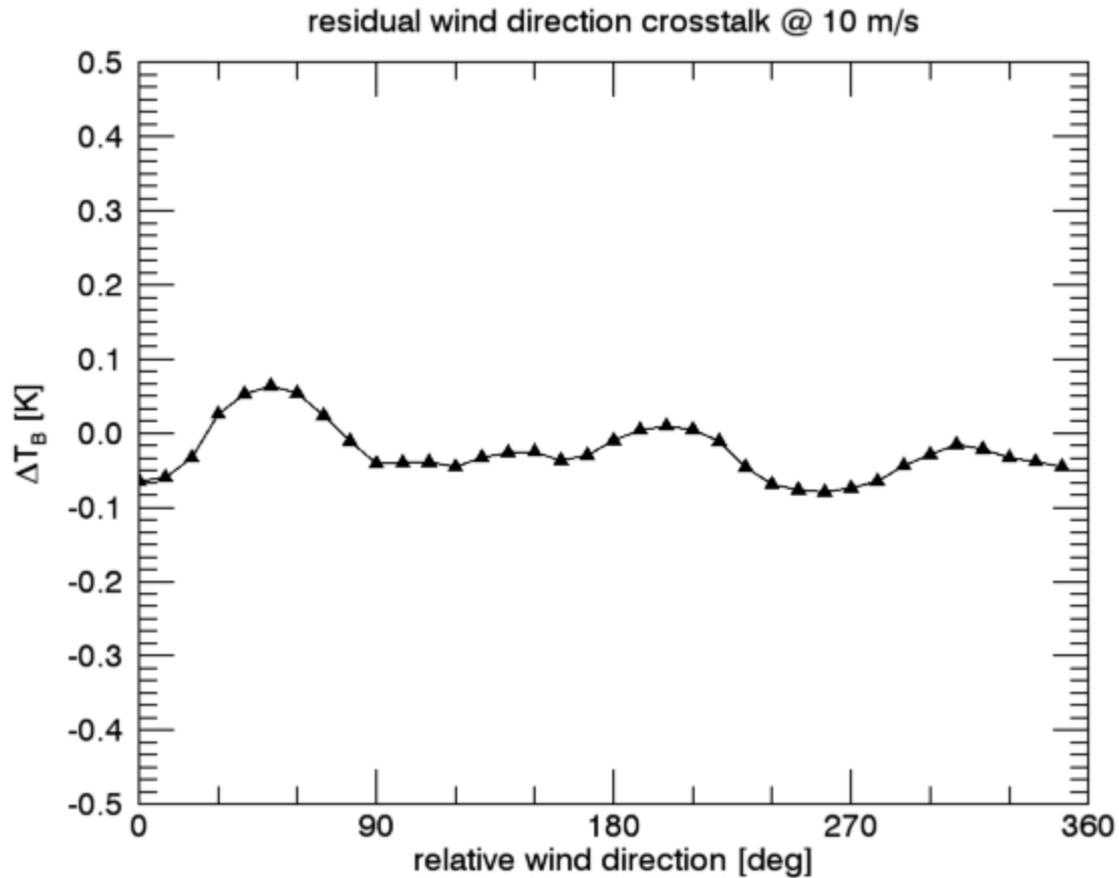
very small - none

Wind Speed: Aquarius versus NCEP

NCEP and Aquarius wind speed differences are very small during times of peak biases.

Wind Speed: Aquarius versus WindSat
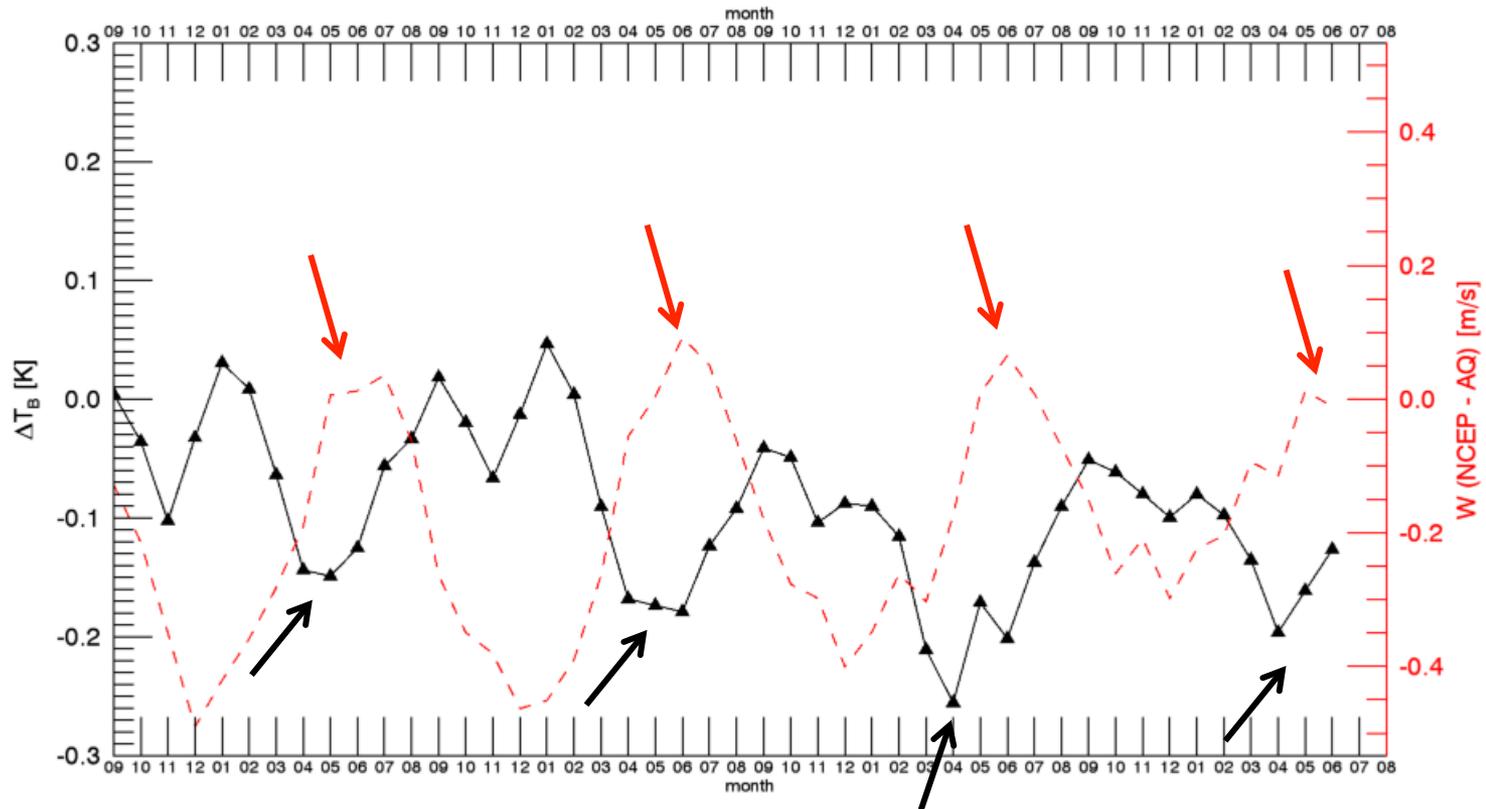
WindSat and Aquarius wind speed differences are very small during times of peak biases.

HHH – WindSat
Inter-Beam Differences

V-pol/H-pol

An error in input wind speed would affect H-pol stronger than V-pol.

Ancillary SST Input

ΔSST (h-pol) [K] horn 3
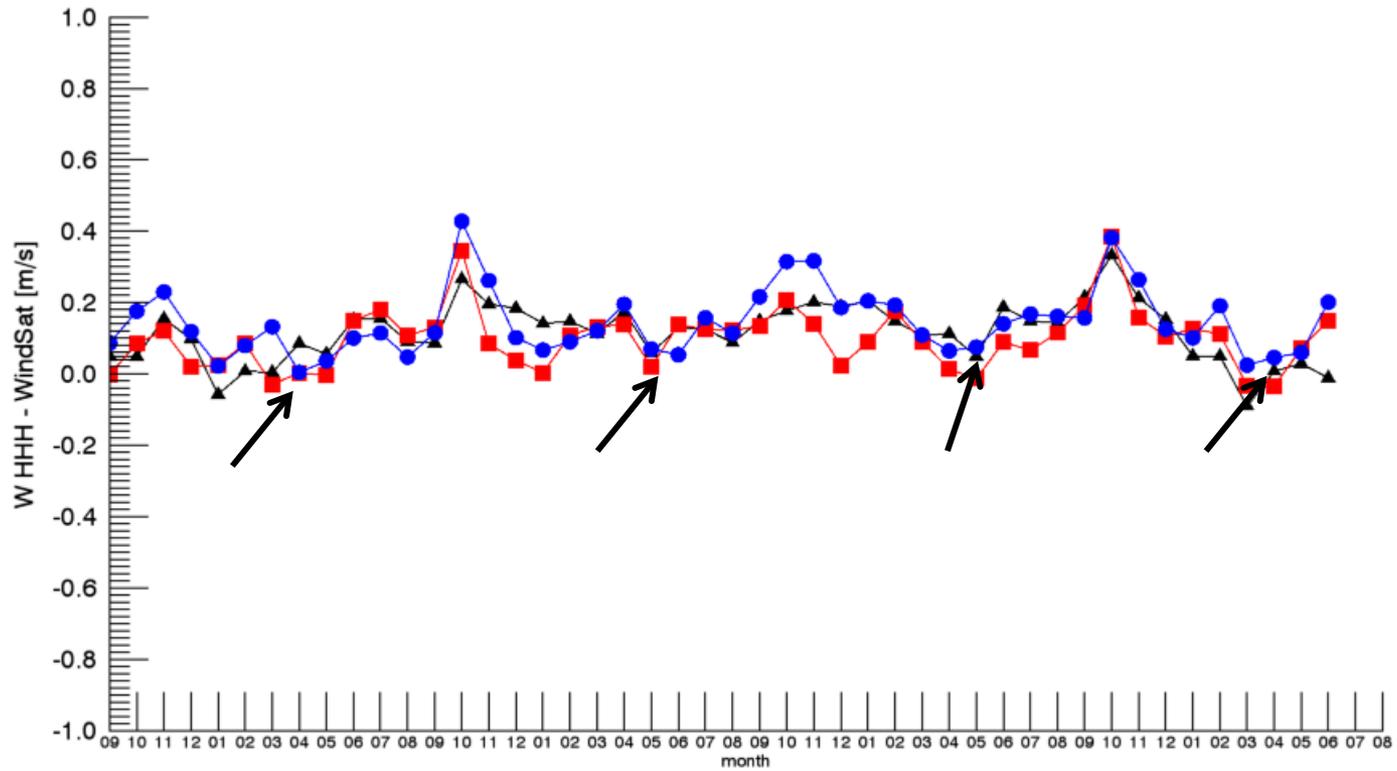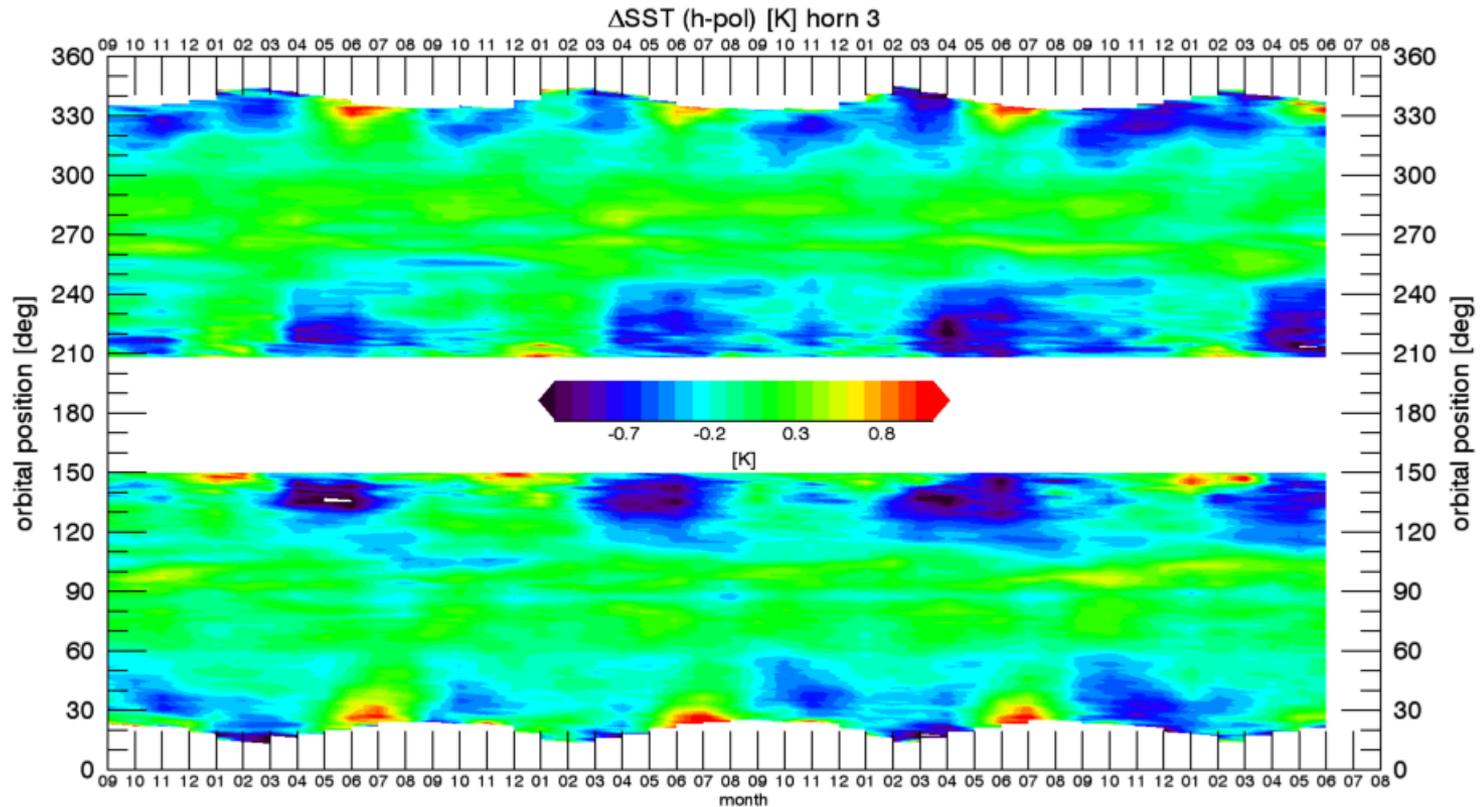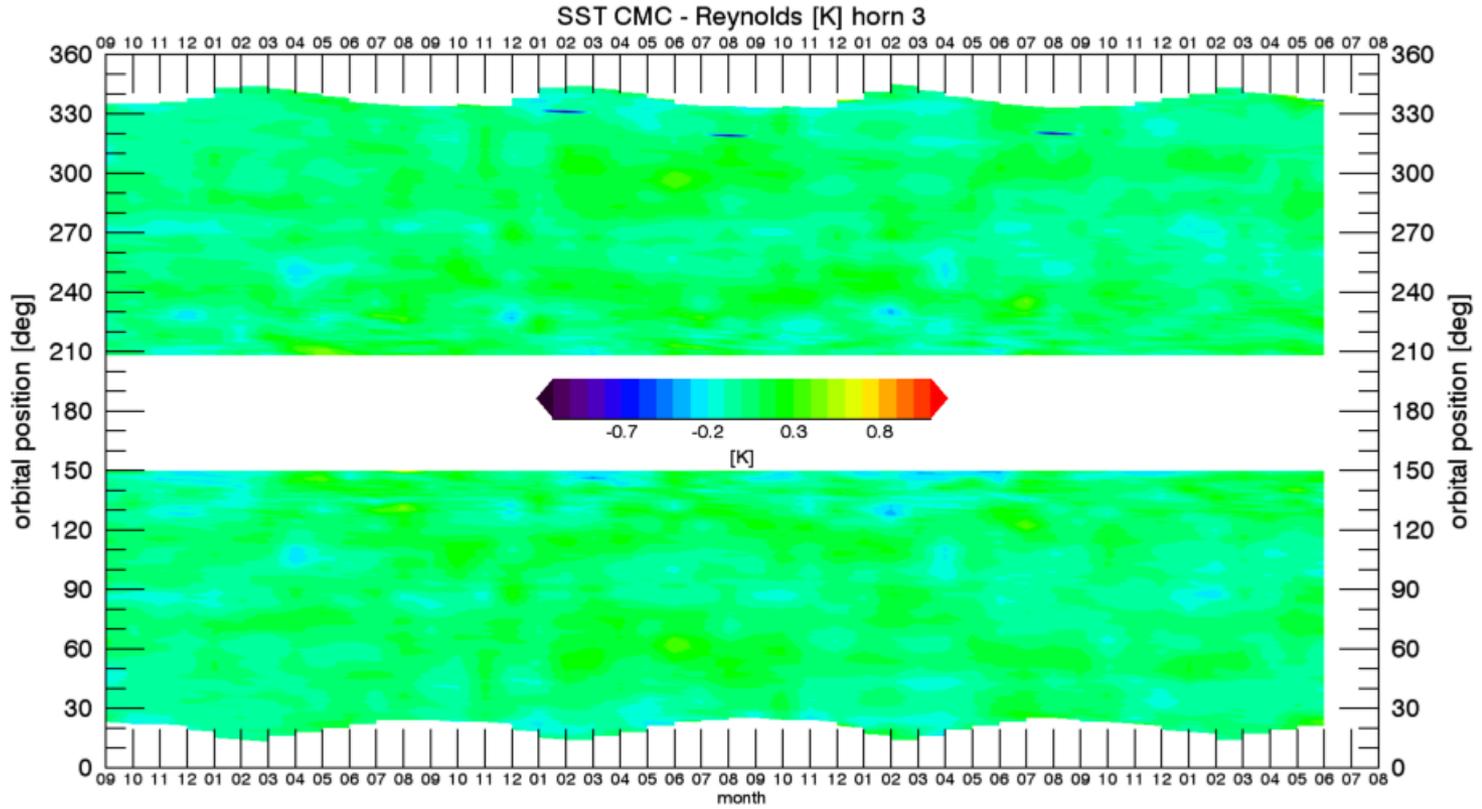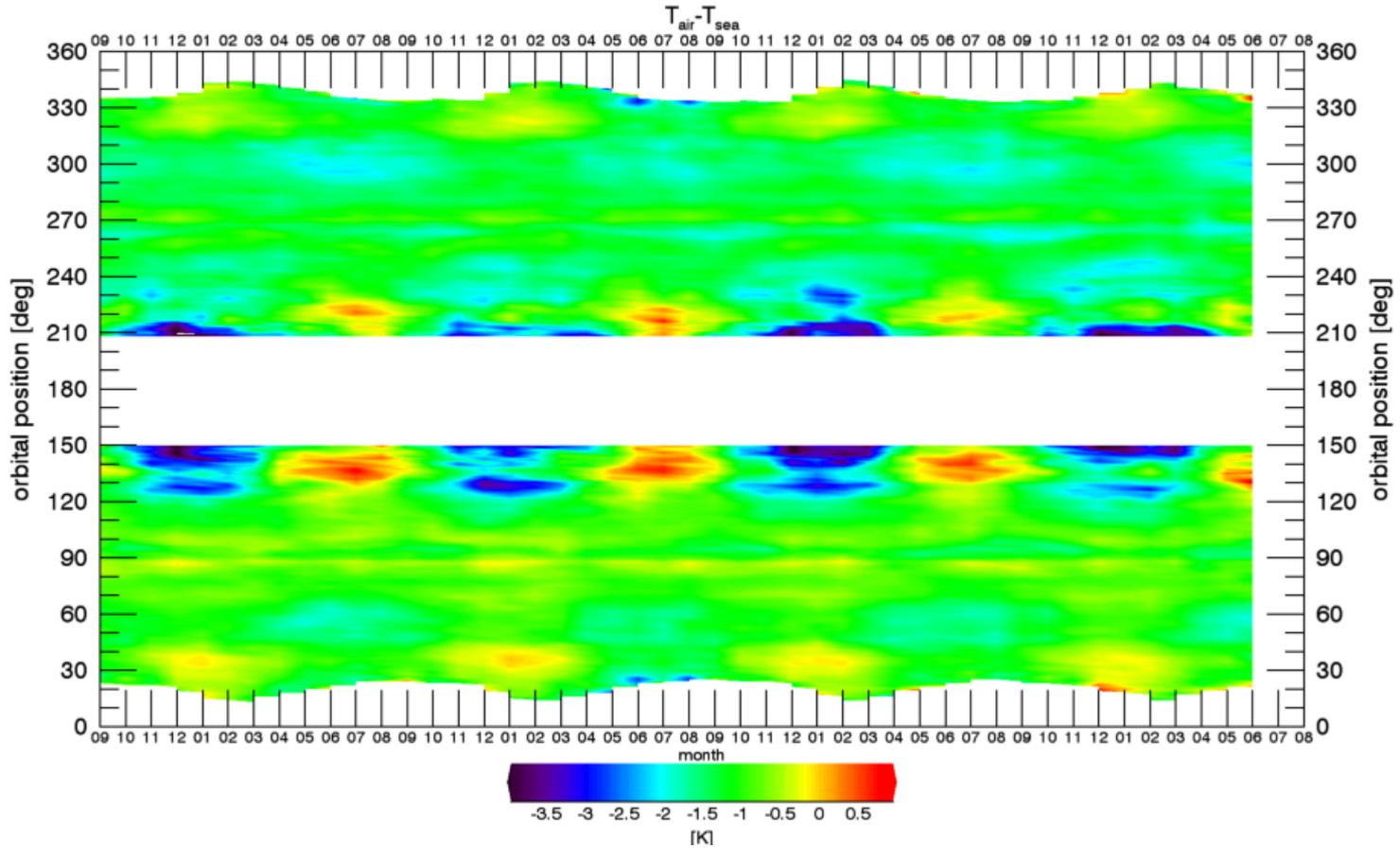
The figure shows the change in SST (from Reynolds) that one would need to explain the biases. The SST error would need to be very large.
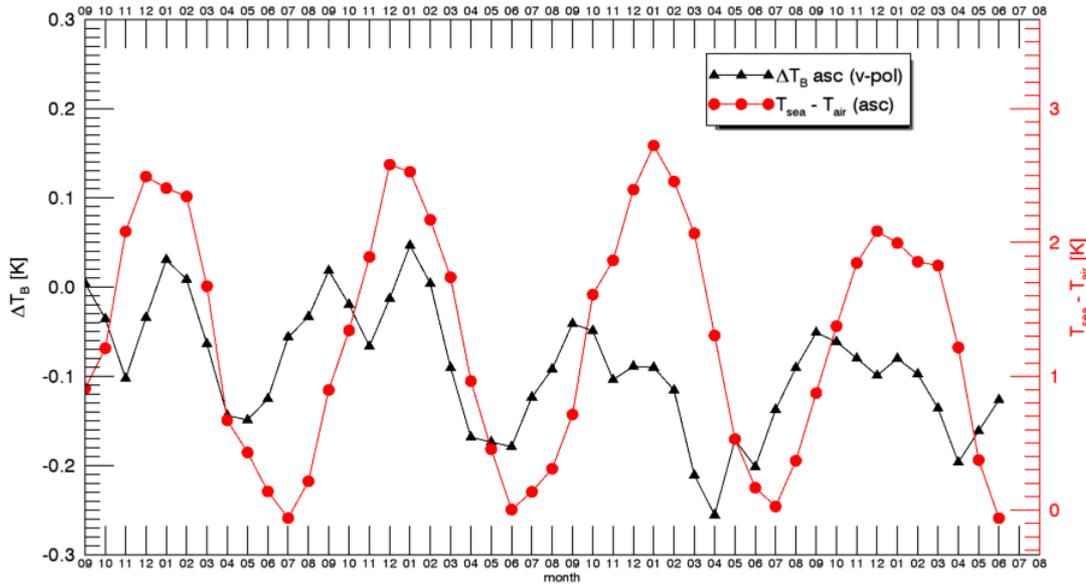Not observed at SMAP.

CMC – Reynolds SST

SST CMC - Reynolds [K] horn 3

shows nothing like that

Air – Sea Temperature

A close look:

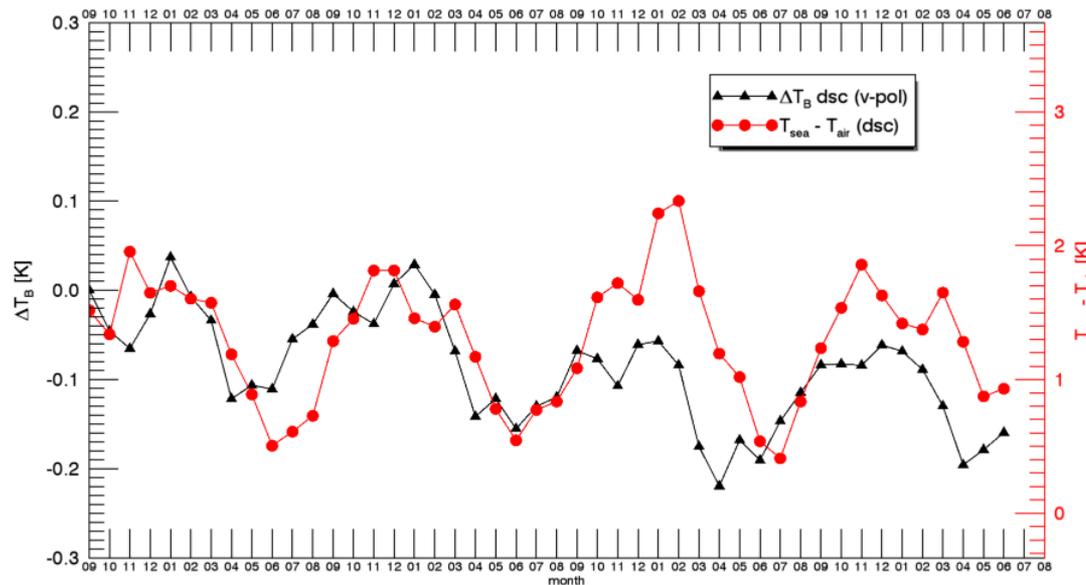Aquarius TB bias in N hemisphere peaks during MAM.

Tair – Tsea peaks in MJA.

Aquarius TB bias in N hemisphere has little-no asc/dsc difference.

Tair – Tsea has significant diurnal signal.

Residual Dependence on T_{air} - T_{sea}
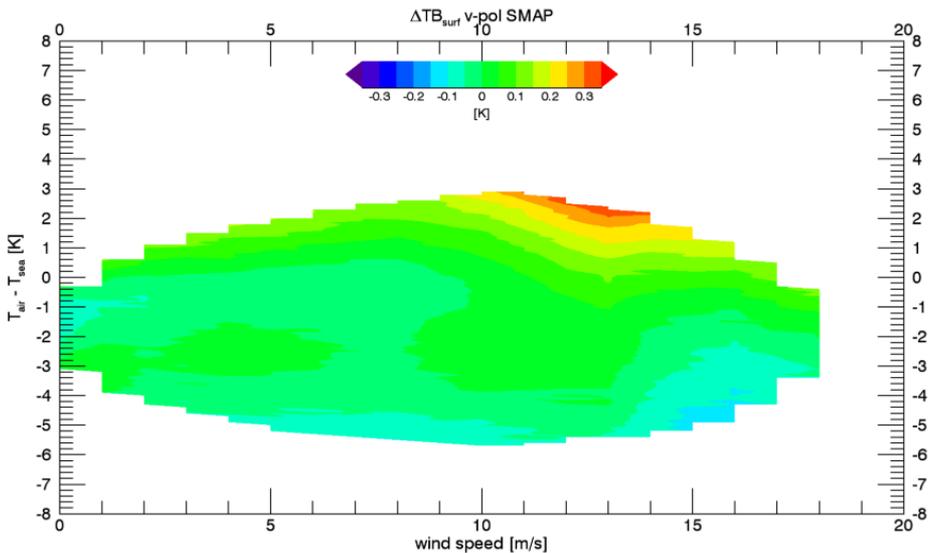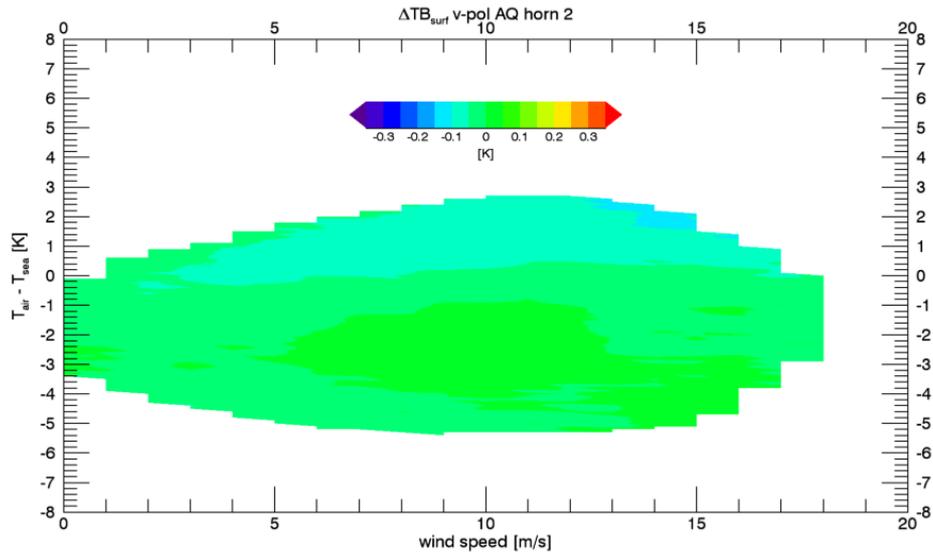
There is none in either Aquarius nor in SMAP.

Using scatterometer (HHH) and WindSat in surface roughness correction takes care of it.

(Same applies to SWH).

ΔTB$_{surf}$ v-pol AQ horn 1

ΔTB$_{surf}$ v-pol AQ horn 2

ΔTB$_{surf}$ v-pol AQ horn 3

ΔTB$_{surf}$ v-pol SMAP

V3.0 Aq-HYCOM bias adj. SSS vs Air-Sea dT for 25N to 60 N

Strong correlation between salinity bias and air-sea temperature difference:
~0.1 psu/C

focus!

from S. Brown
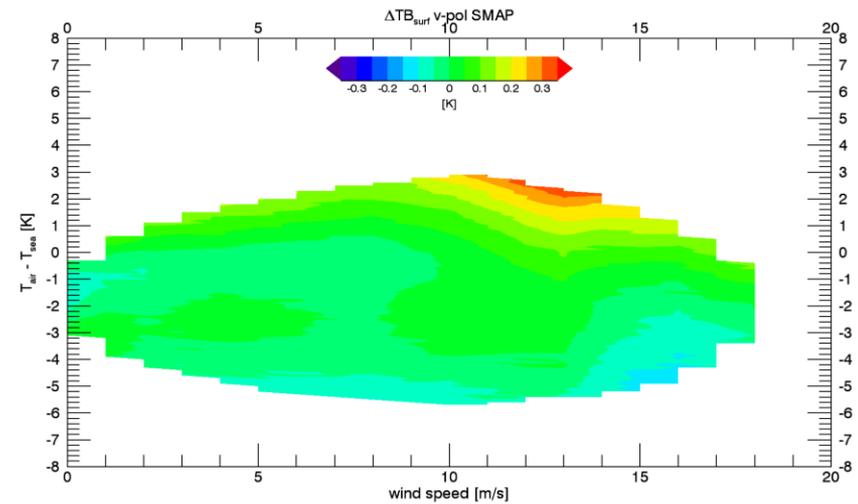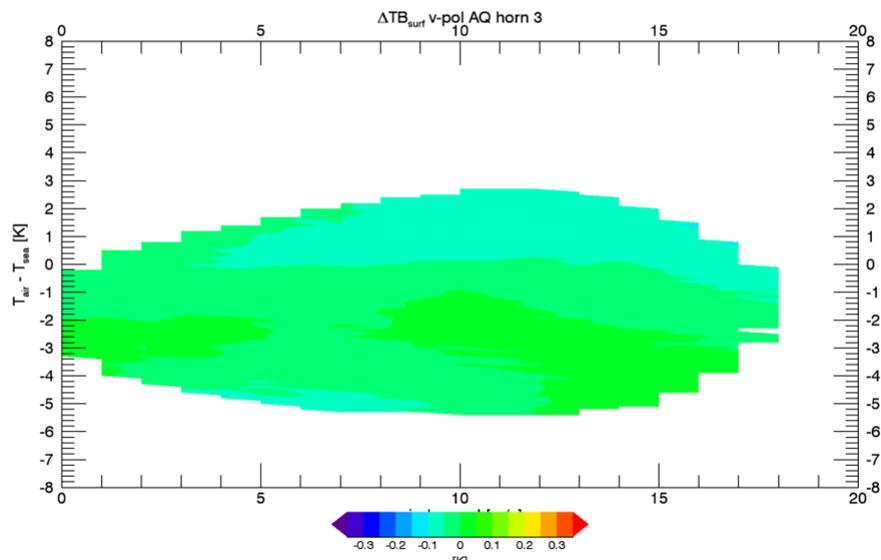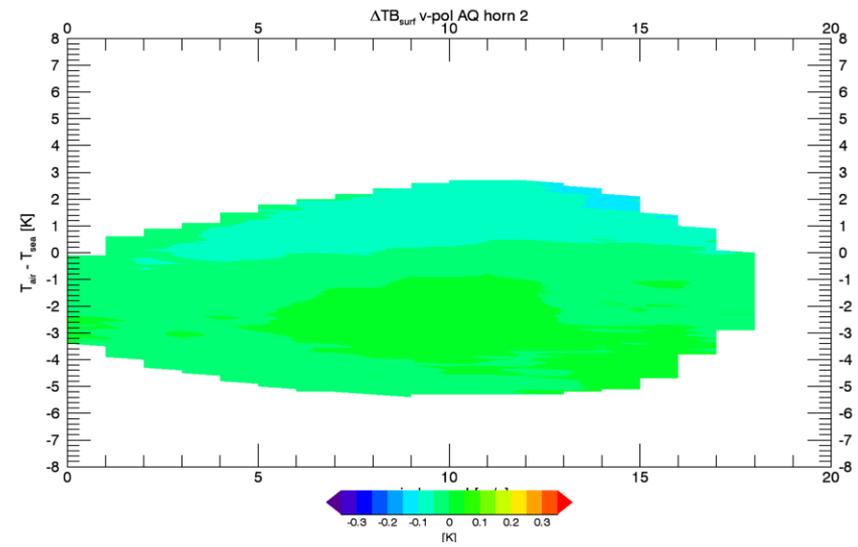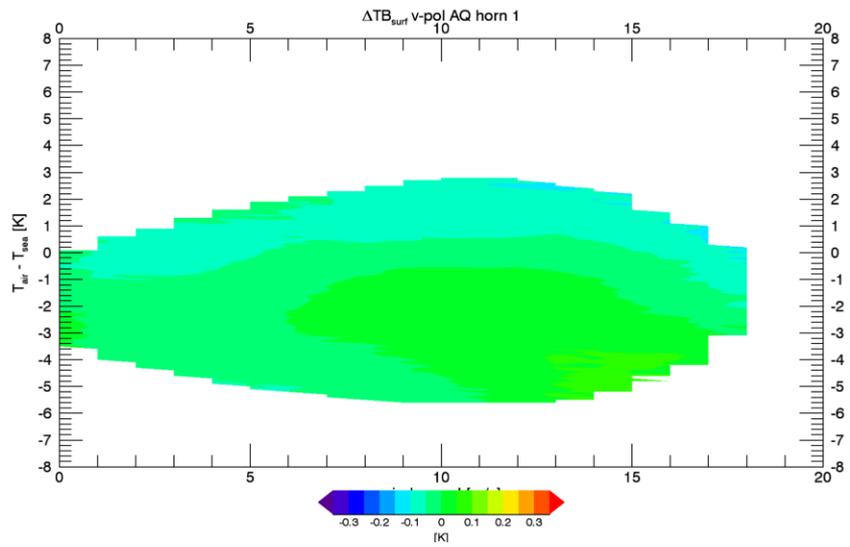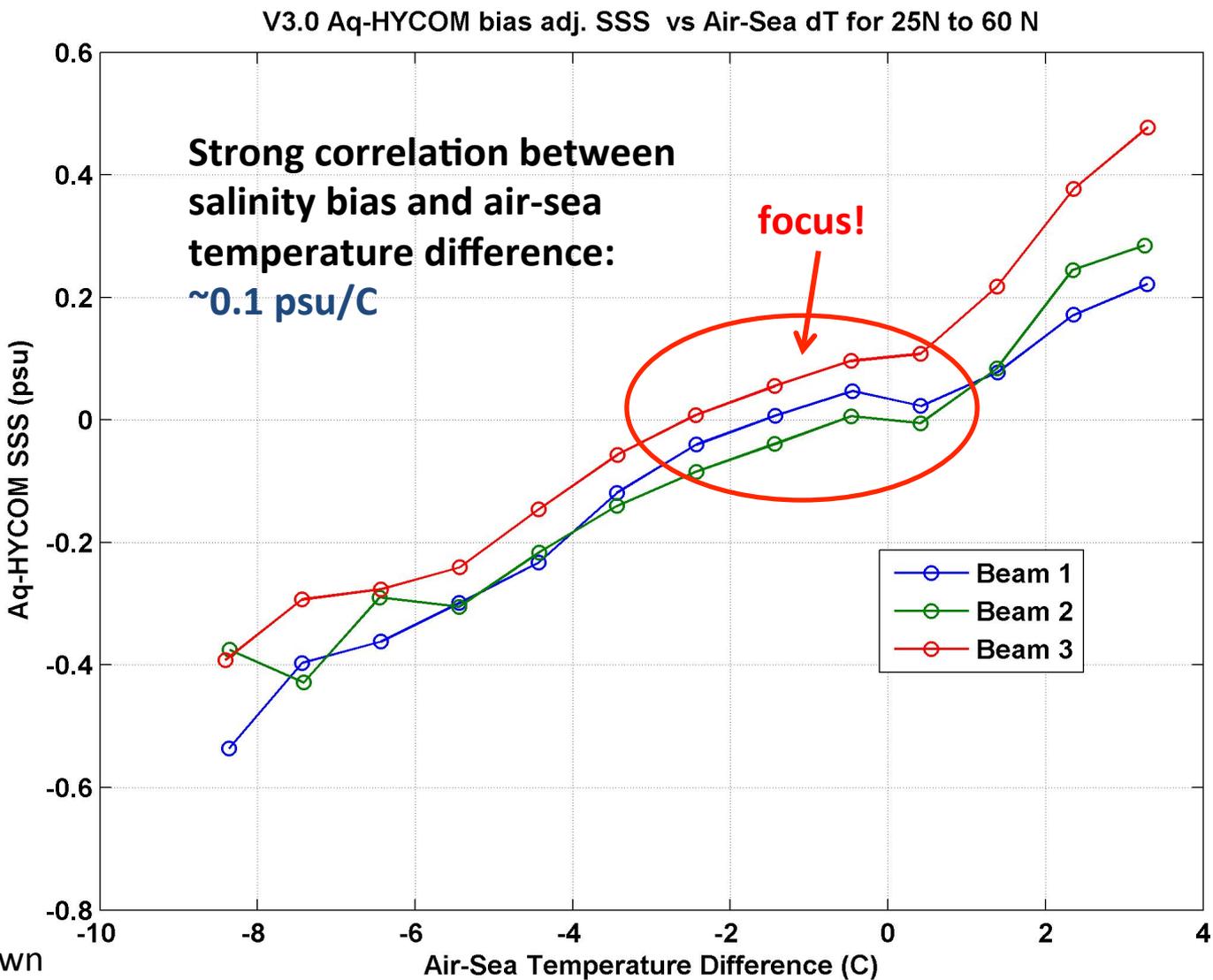
Summary/Conclusions (1)

- Analysis of the N latitude biases in Aquarius V4.
  - Biggest problem in V4.
- Main pattern
  - horn1 ≈ horn3 > horn 2
  - Exceeds 0.25K (0.5 psu) in horn 1 and horn 3.
  - V-pol > H-pol
  - peak in MAM
  - ascending ≈ descending
  - Inter-annual variability: Weak in 2012. Strongest in 2014.

Summary/Conclusions (2)

- Not seen in SMAP (so far) and SMOS
- There are no indications that these biases can be reasonably explained within the geophysical model
 - Galaxy, sun, …
  - Surface roughness model: SST
  - Surface roughness model: wind speed
  - Surface roughness model: wind direction
  - Input wind speed
  - Input SST
  - Atmosphere
  - Air – sea temperature
  - SWH

Summary/Conclusions (3)

- Increasingly likely that it is Aquarius specific and instrument related
  - like it was the case with the non-linear IU coupling (only seen at Aquarius, absent at SMAP).
- Not clear if it is related to spurious annual oscillation in V4.1
- What to do about it in V5.0 ?!
- Bias in S hemisphere
  - Potentially galaxy, high winds

TB V-pol – H-pol (2nd Stokes)
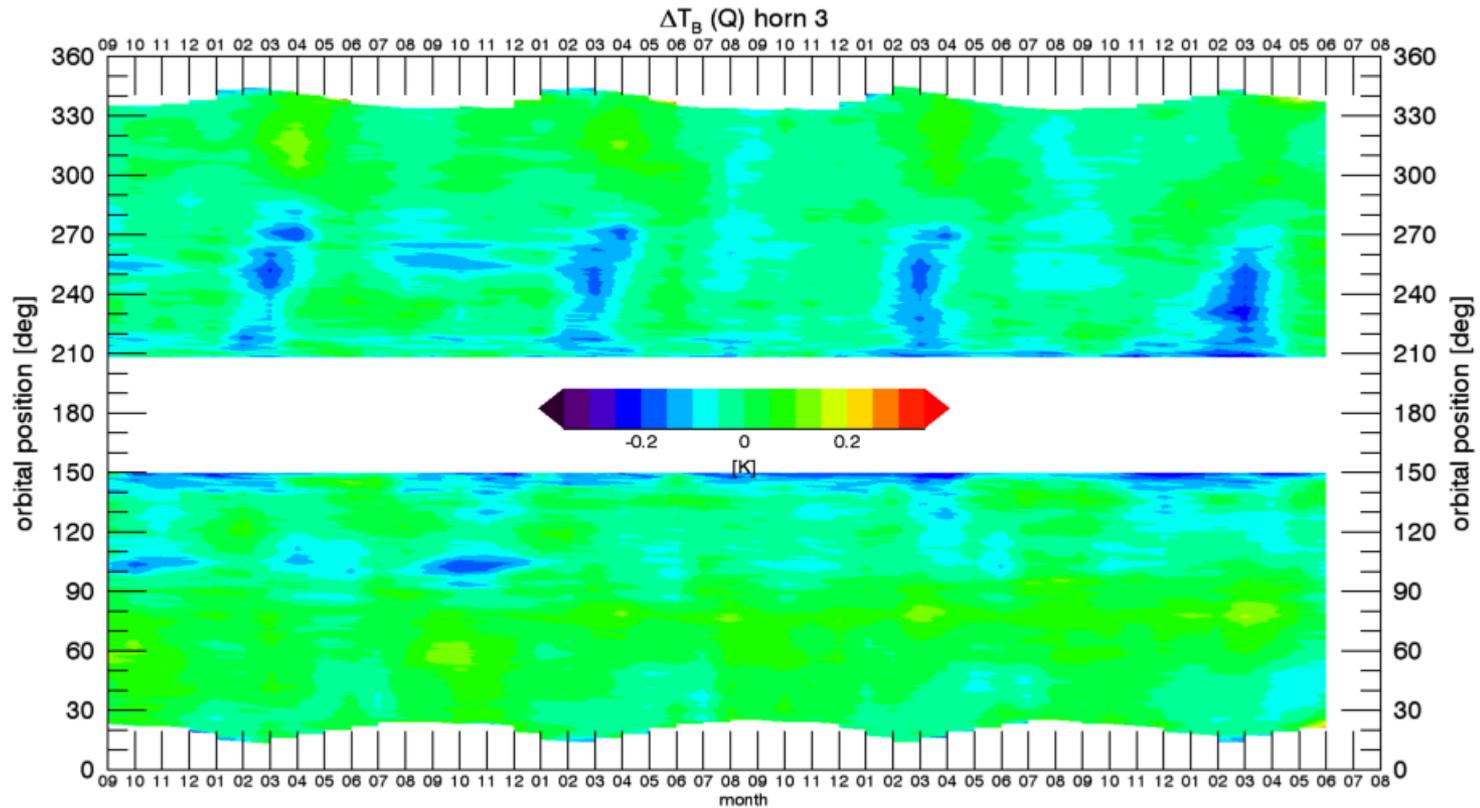
asc/dsc bias in 2nd Stokes
issue with 3rd Stokes?

Backup Slides

ΔT_B V-pol [K] horn 3

TA gal ref [K] horn 3