



# The RSS/NASA SMAP Salinity Version 3 Release

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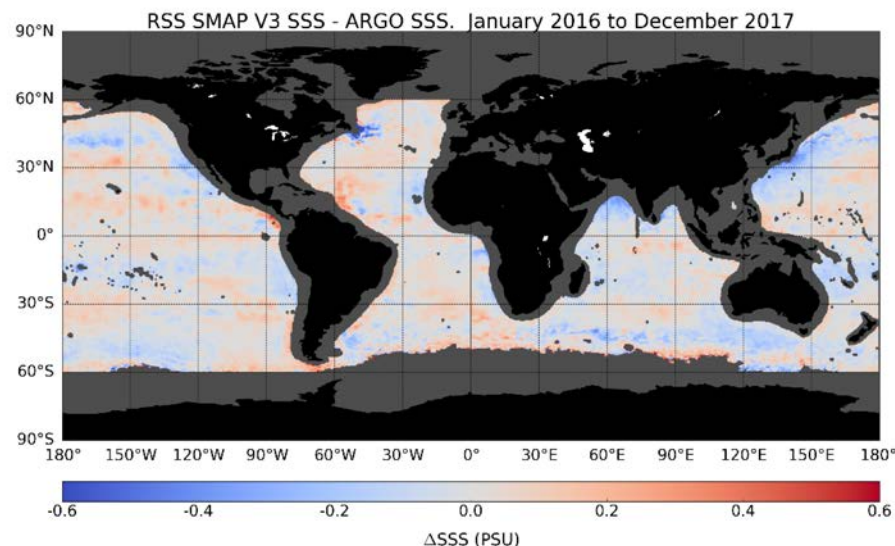
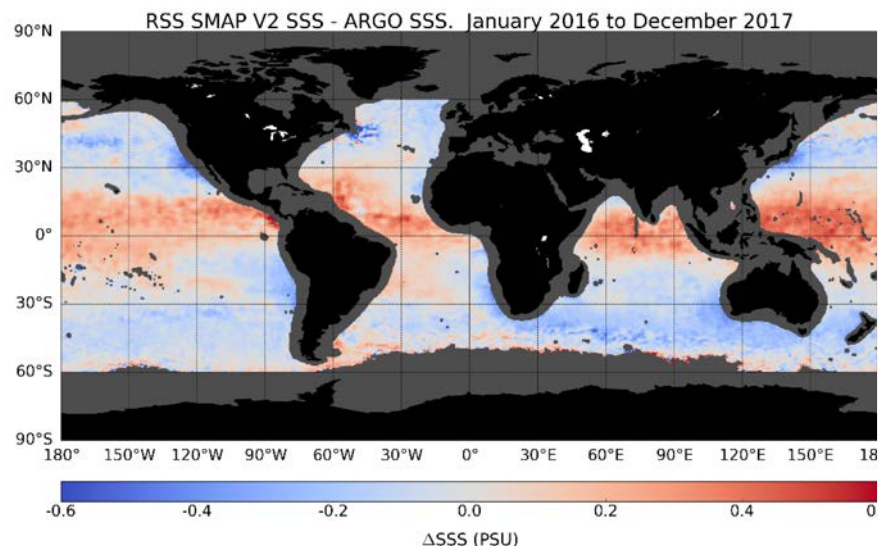
# Outline

- Overview of RSS SMAP V3 Release
  - Improvements over V2
  - Performance
  - Data Distribution
- Spatial Sampling and Noise Figures
- V3 Retrieval Algorithm
  - Geophysical Model
  - Ancillary Wind Speed
  - Emissive SMAP Reflector
  - Land Correction
- Aquarius – SMAP Continuity (NASA SCPS)

# SMAP – ARGO (Scripps) rain filtered, open ocean

V2.0

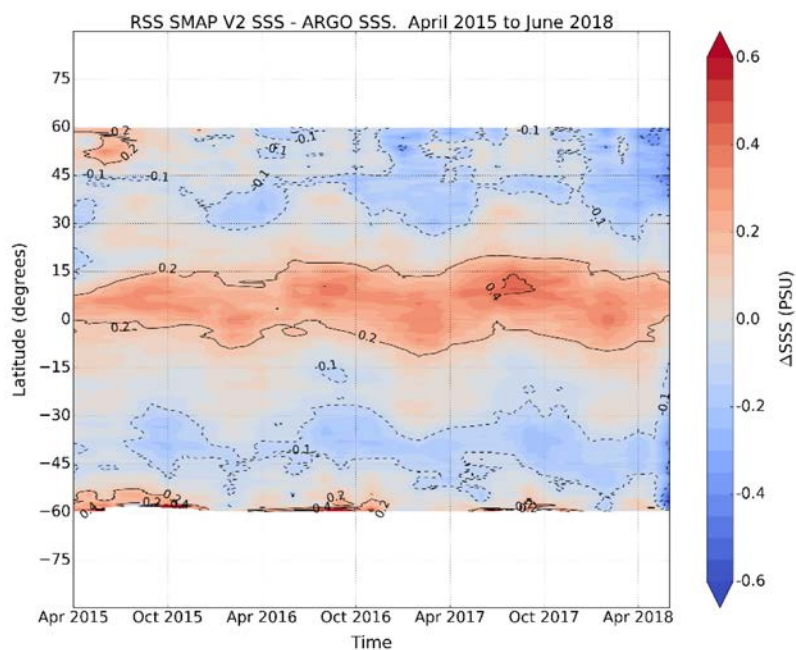
V3.0



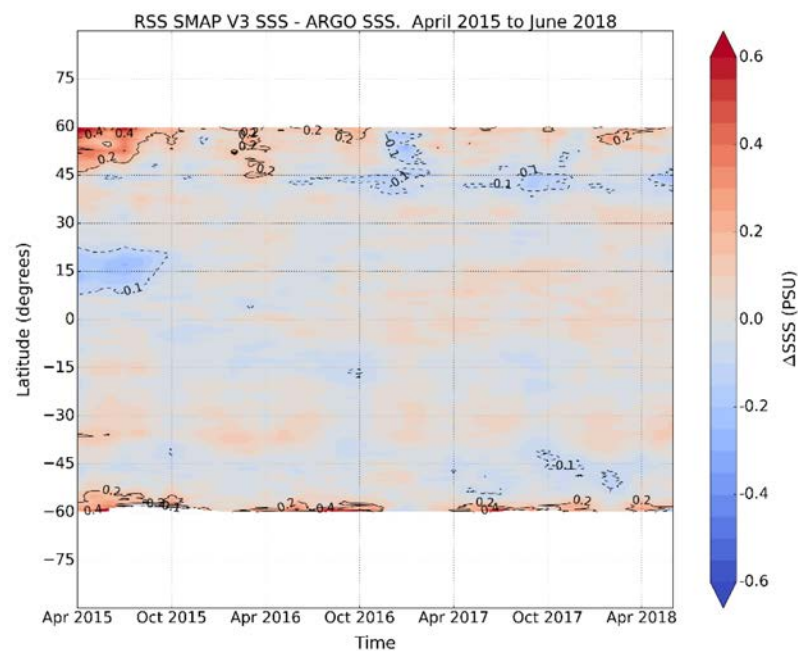
V3 removes spurious salty biases in the tropics that were prevalent in V2.

# SMAP – ARGO (Scripps) rain filtered, open ocean

V2.0



V3.0



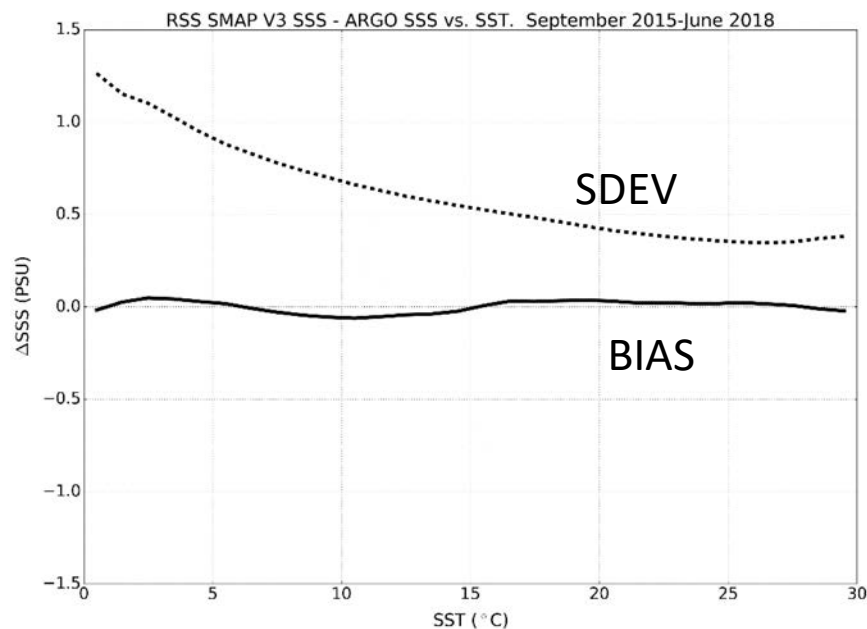
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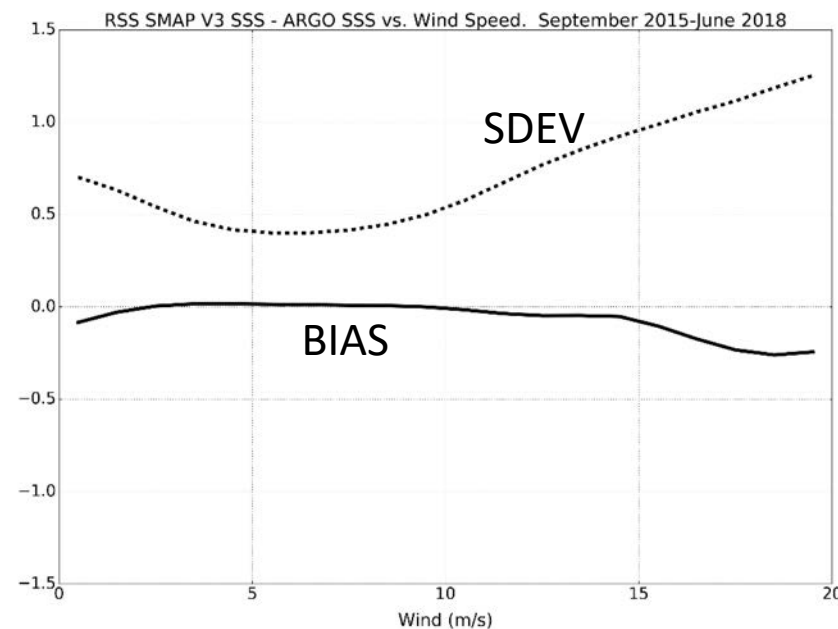
# RSS SMAP V3.0 Level 2 – ARGO

## Open Ocean. Binned v SST and Wind Speed

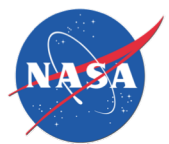
### VS. SST



### VS. WIND SPEED



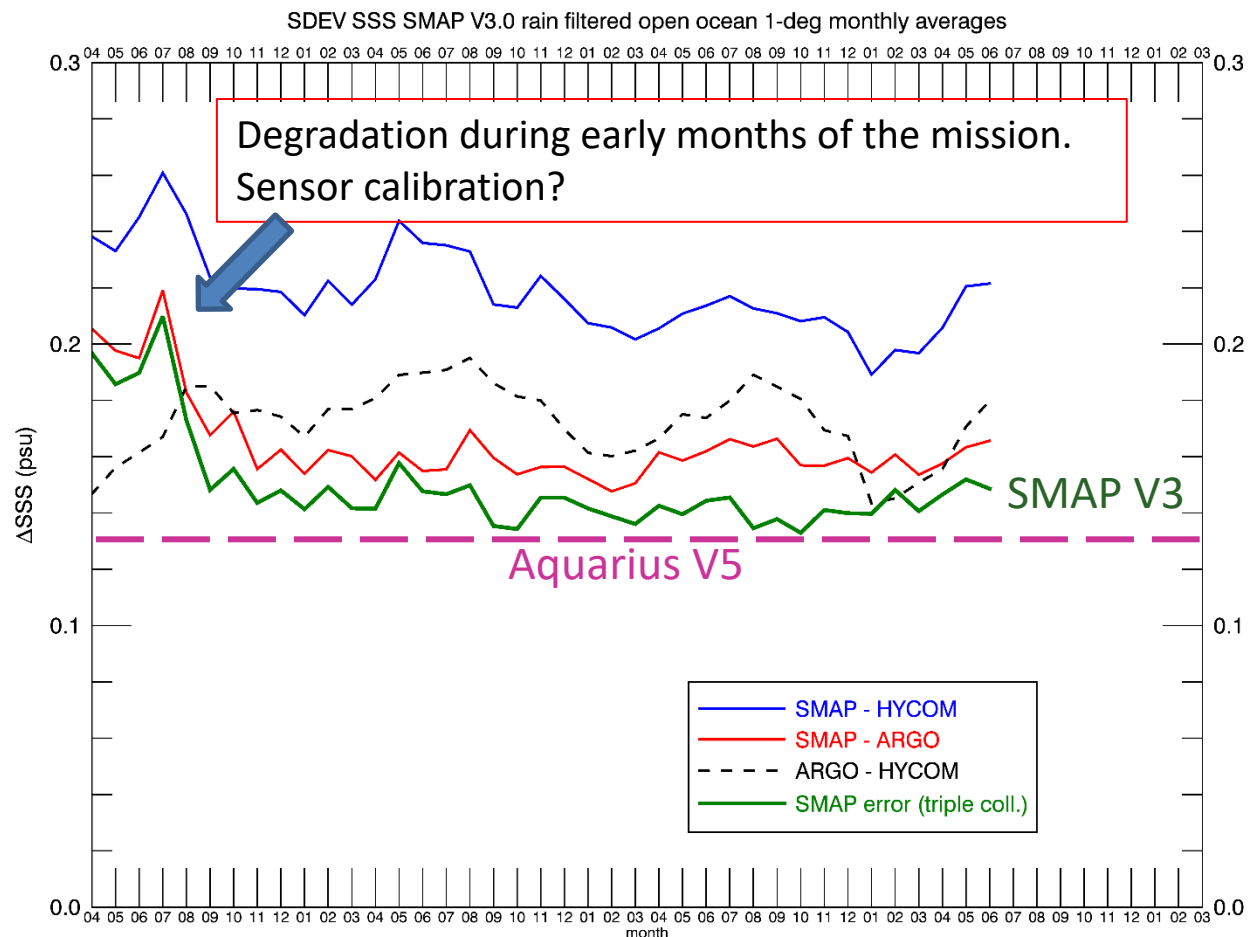
No biases over whole SST range and wind speed range (< 15 m/s).  
Even in cold water.  
Consistency of GMF (dielectric model, wind emissivity model).



# SMAP (V3) – ARGO – HYCOM

Triple Collocation. 1-deg Monthly Averages.

SMAP V3 reaches quality of Aquarius V5





# NASA/RSS SMAP V3 SSS

## Data Products

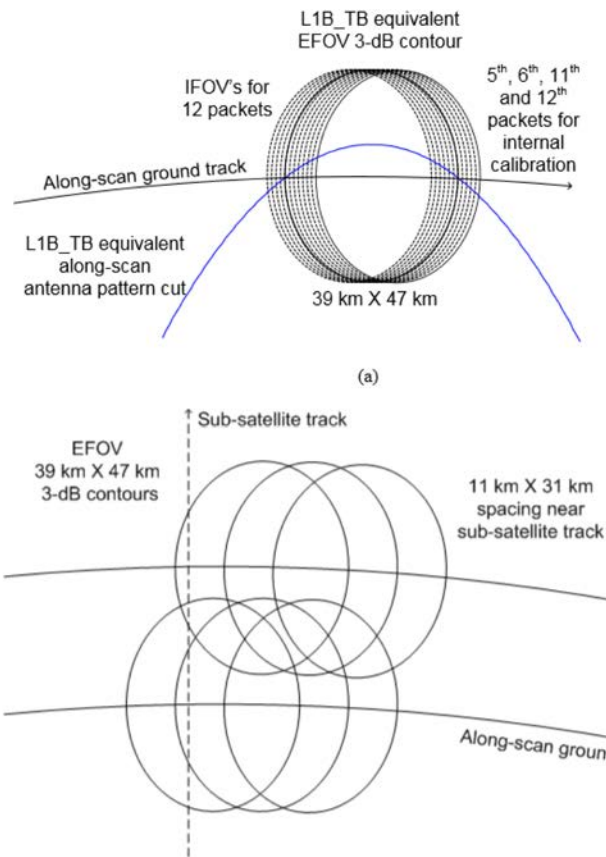
40 km	70 km standard product	content	latency
L2C	L2C	$\frac{1}{4}$ deg swath grid SSS + all essential orbital and pointing parameters, all TA, TB, ancillary fields, all corrections, forward model (TA expected)	$\sim 1$ day
L3 8-day running	L3 8-day running	$\frac{1}{4}$ deg map SSS + land fraction, sea-ice fraction, SST	$\sim 1$ week after center day
L3 monthly	L3 monthly	$\frac{1}{4}$ deg map SSS + land fraction, sea-ice fraction, SST	$\sim 1$ week after end of the month
	L3 monthly rain filtered (RF) @ RSS only	$\frac{1}{4}$ deg map SSS + land fraction, sea-ice fraction, SST	

**RSS** [www.remss.com/missions/smap/salinity](http://www.remss.com/missions/smap/salinity)

**NASA PO.DAAC** <ftp://podaac-ftp.jpl.nasa.gov/allData/smap/>

# SMAP

## Footprint – Sampling – Noise



L1B sampling time:

- Aquarius: 1.44 sec (pre-averaged)
- SMAP: ~ 14 msec (EFOV)

**results in 10 times larger noise in SMAP L1B**  
strongly oversampled (overlap between EFOV)

**Noise figure for SMAP L1B EFOV is ~ 2 psu.**

- All random noise.
- **Useless** for most oceanic applications.

**RSS/NASA SMAP L2 processing resamples SMAP L1B** using Backus-Gilbert type Optimum Interpolation (OI) onto fixed  $\frac{1}{4}$  deg Earth grid.

- Utilize over-sampling in along-scan direction.
- 40 km elliptical (keep original resolution)
- **70 km circular**

from SMAP L1B ATBD (J. Piepmeier et al.)





# SMAP

## Approximate Noise Figures + Spatial Resolution for Salinity Product

Product	Spatial Resolution	Noise	Noise Type
SMAP L1B	40 km	~ 2 psu	random (NEDT)
RSS SMAP L2	40 km OI	~ 0.9 psu	random (NEDT)
<b>RSS SMAP L2</b>	<b>70 km OI</b>	<b>0.5 psu</b>	<b>random (NEDT) + systematic (GMF)</b>
<b>RSS SMAP L3 monthly</b>	<b>70 km</b>	<b>0.15 psu</b>	<b>systematic (GMF)</b>

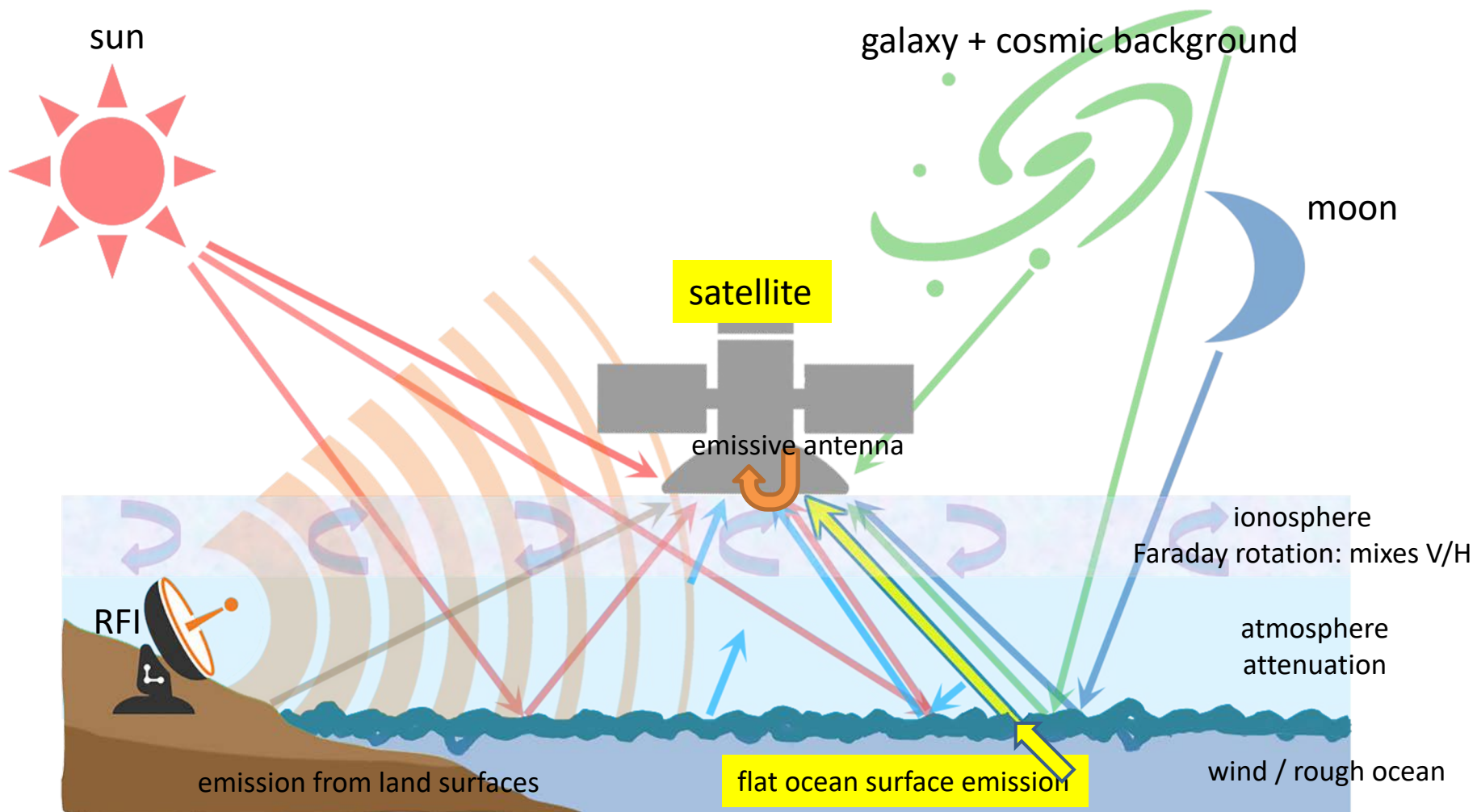
The OI results in a significant noise reduction of an otherwise extremely noisy product.

For most open ocean applications, the 70-km products are the best to use as they have significantly lower noise.

# Salinity Retrieval

Large spurious signals

Only about 10% of the received signal comes from SSS





# Salinity Retrieval

## Geophysical Model (GMF)

- Requires very accurate GMF to remove all these signals
  - 0.2 psu salinity accuracy requires 0.1 Kelvin radiometric accuracy
  - worse in cold water
- We cannot use the Level 1 brightness SMAP temperatures from NSIDC for retrieving salinity
  - Very good RFI filter
  - Recent Version 4 L1B SMAP TB improved from prior versions
  - radiometric accuracy  $\sim 1$  Kelvin
  - Does not fulfill the salinity accuracy requirements: galaxy, atmosphere, emissive antenna, cross-pol, wind roughness
- RSS SMAP ocean products
  - start from TA (antenna temperature) before correcting for emissive antenna
  - L1B TB processing
- We cannot use canned radiative transfer models (community RTM, ...) to retrieve salinity
  - TB assimilation



# SMAP V3 Salinity Retrieval Algorithm

output of all intermediate steps included in L2C files

**SMAP L1B Antenna Temperatures (TA)**

RFI filtered.

Before correcting emissive antenna.

**L2A Pre-Processing**

OI onto fixed  $\frac{1}{4}$  deg Earth grid  
(40 km / 70 km)

remove antenna emission

TA

remove space contributions:  
galaxy, sun, moon, cold space

Earth TA

remove antenna x-pol

Top of the Ionosphere (TOI)  
Brightness Temperatures

**Ocean Target Calibration**  
**SSS SMAP = SSS HYCOM**  
global, multi-day scale

correct for Faraday Rotation  
in Earth Ionosphere (3<sup>rd</sup> Stokes)

**SMAP L2  
Salinity**

find salinity for which emissivity  
of Meissner-Wentz 2012 dielectric  
model matches specular TB

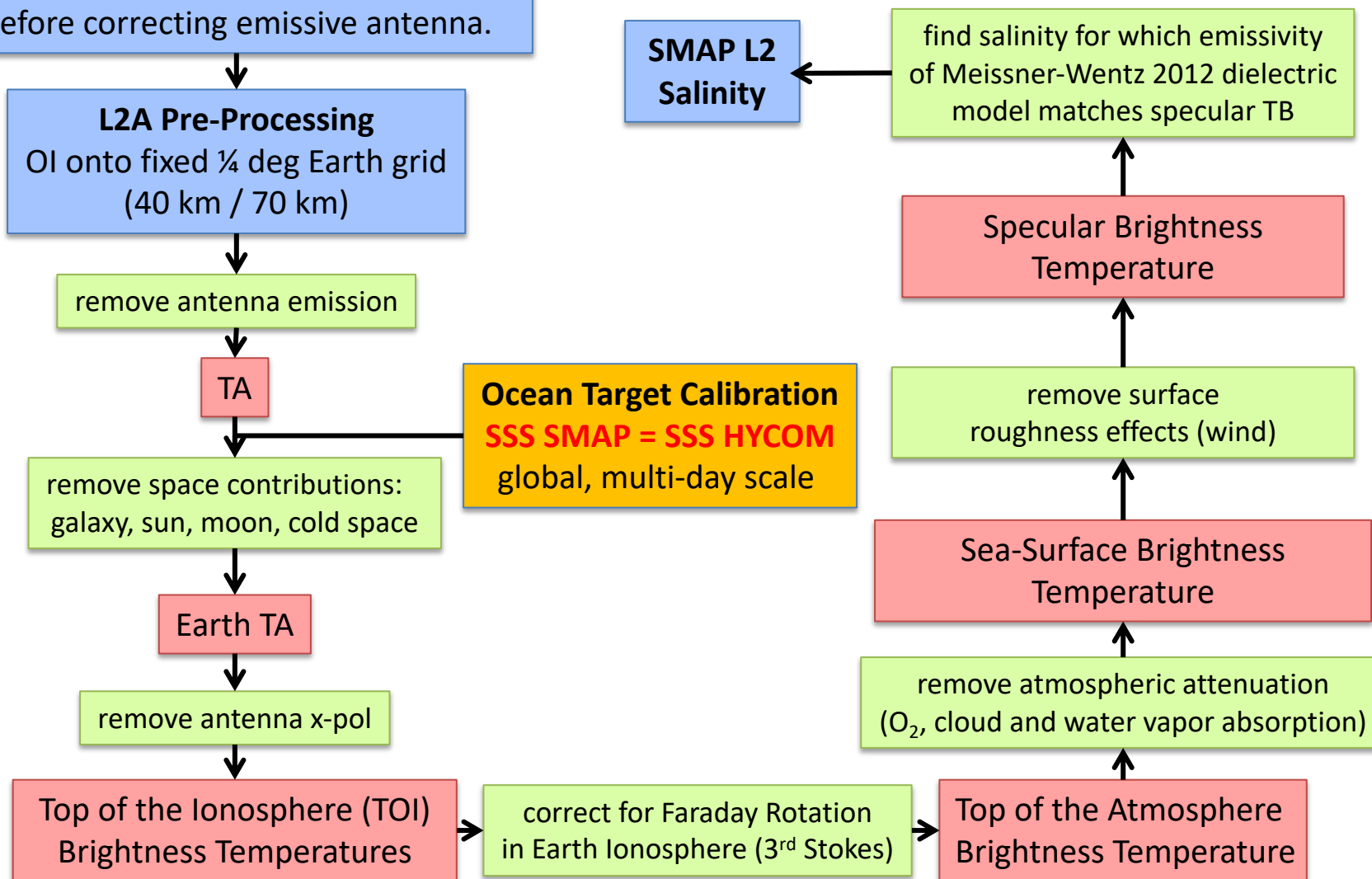
Specular Brightness  
Temperature

remove surface  
roughness effects (wind)

Sea-Surface Brightness  
Temperature

remove atmospheric attenuation  
(O<sub>2</sub>, cloud and water vapor absorption)

Top of the Atmosphere  
Brightness Temperature





# SMAP V3 Salinity Retrieval Algorithm

## Major Changes from V2 Release

### 1. GMF: Consistent with Aquarius V5 release.

- i. Meissner Wentz dielectric model.
- ii. Liebe oxygen absorption (important to mitigate biases in Aquarius V5).
- iii. Wind emissivity. (Caveat: different ancillary wind speeds).
- iv. Galaxy (based on SMAP fore – aft, already implemented in V2).
- v. **Use ARGO as reference SSS to derive GMF.**  
**Avoid biases between ARGO/HYCOM.**

### 2. CCMP ancillary wind speed + direction.

### 3. Improved land correction.

### 4. Input SMAP Version 4 L1B TA (latest calibration).

### 5. Include IMERG rain rate (L2).

- Resampled to SMAP resolution: 40 km/70 km.
- Used for rain flagging (comparison with ARGO).
- Used for atmospheric cloud water correction (instead of NCEP cloud water).



# Surface Roughness Correction

## CCMP (Cross Calibrated Multi-Platform) Winds

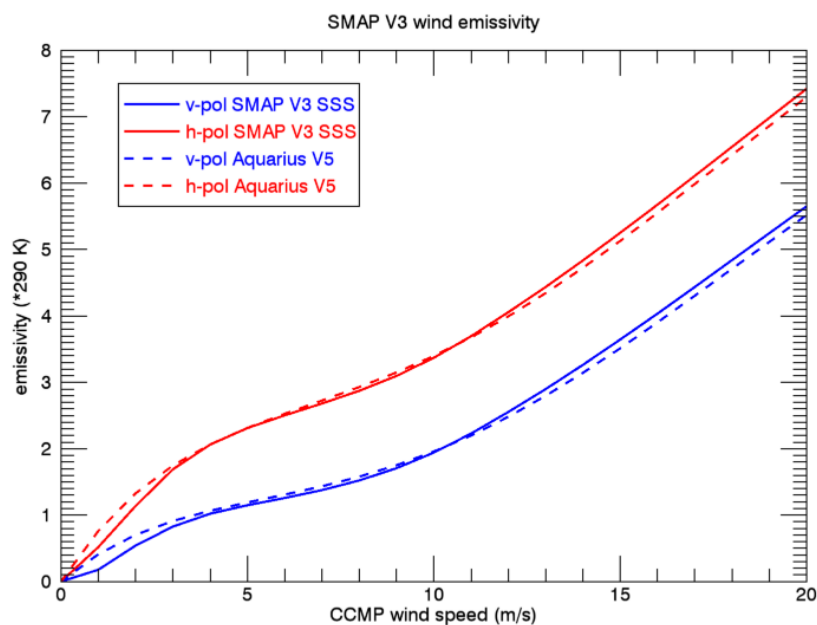
- **Surface roughness is the largest error source**
  - **0.3 m/s error in wind speed makes 0.2 psu error in SSS**
- SMAP lost its scatterometer.
  - Information in radiometer channels is not sufficient to retrieve salinity and wind speed simultaneously.
  - SMAP V/H response to increase in wind is similar to decrease in salinity.
  - **Need external wind speed in surface roughness correction.**
- **CCMP for SMAP V3:**
  - Blended Level 4 wind vector product.
  - Variational Assimilation Method (VAM): RSS satellite wind fields + NWP background field.
  - V1: R. Atlas et al. 1996 + 2012, BAMS.
  - V2: produced by C. Mears at RSS.
- **SMAP salinity retrievals use near-real time CCMP V2.0**
  - All RSS radiometers (WindSat, AMSR2, SSMIS, GMI).
  - NCEP GDAS as background field. 6 hourly.  $0.25^\circ$ .  
Space-time interpolated to SMAP.



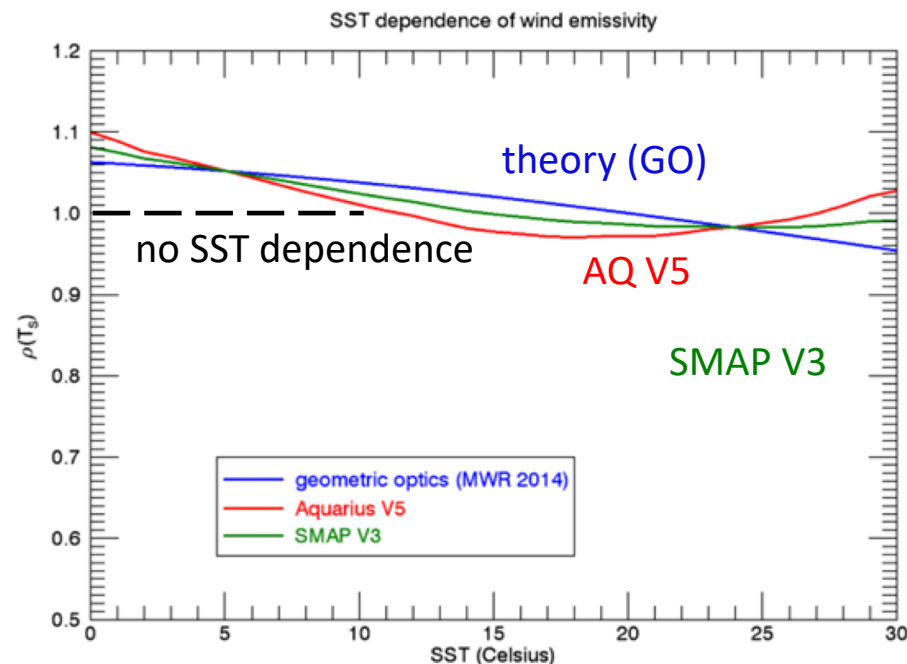
# Wind Emissivity Model

## SMAP V3 – Aquarius V5

Code provided with SMAP V3 release



wind speed dependence



multiply by small SST dependence

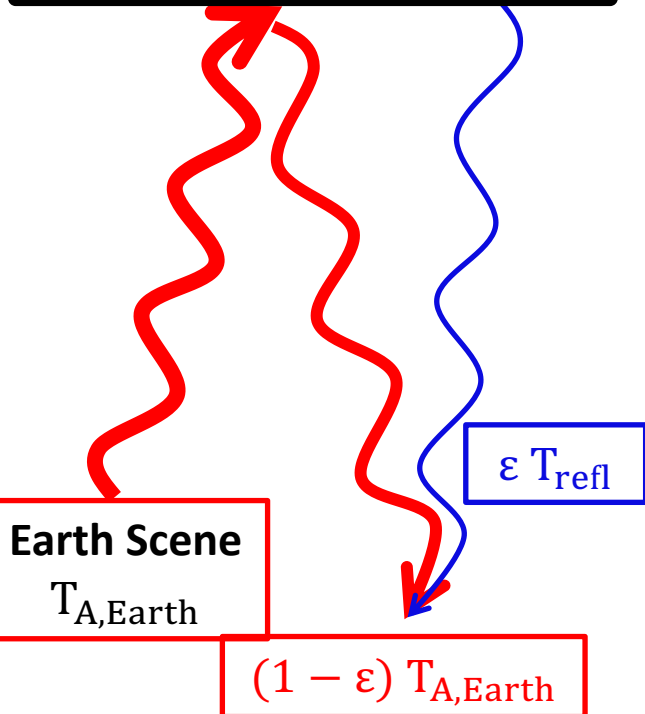
Aquarius V5 uses HHH wind speed (Aquarius scatterometer). SMAP V3 uses CCMP wind speed. There are small systematic differences between these ancillary wind speeds. Rederiving GMF with different wind speed gives slightly different results.

# Emissive Antenna

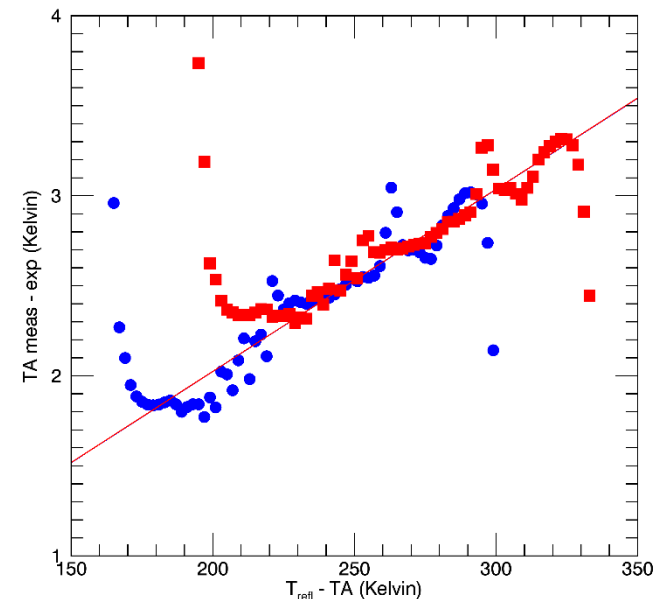
**requires additional calibration effort**

SMAP antenna  
Emissivity  $\varepsilon$

$T_{\text{refl}}$



- The Aquarius antenna was not emissive.
- The SMAP reflector mesh and component of the feedhorn are **lossy (emissive)**.
- $T_A = T_{A,\text{Earth}} + \varepsilon (T_{\text{refl}} - T_{A,\text{Earth}})$
- **Emissivity** about 1%. This is **4 times larger** than pre-launch value.





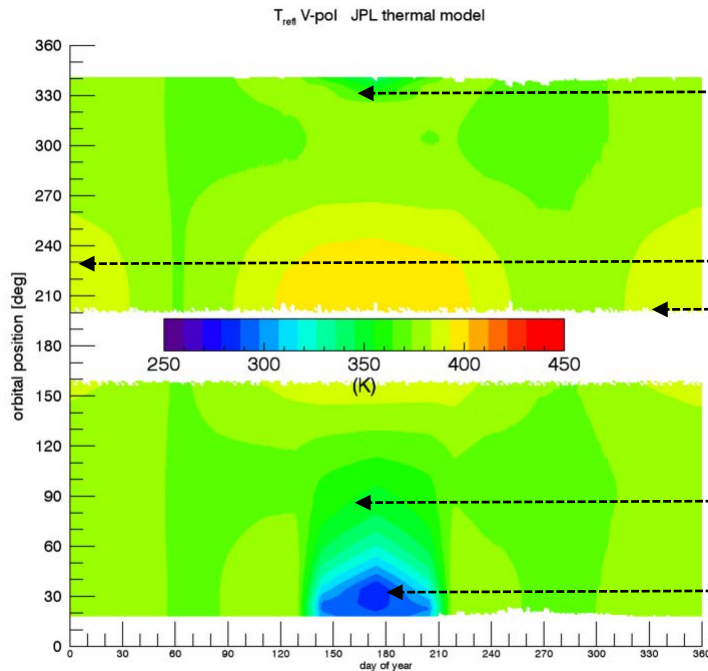
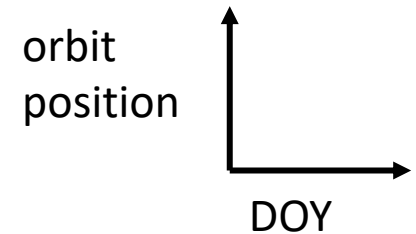


# Reflector Temperature

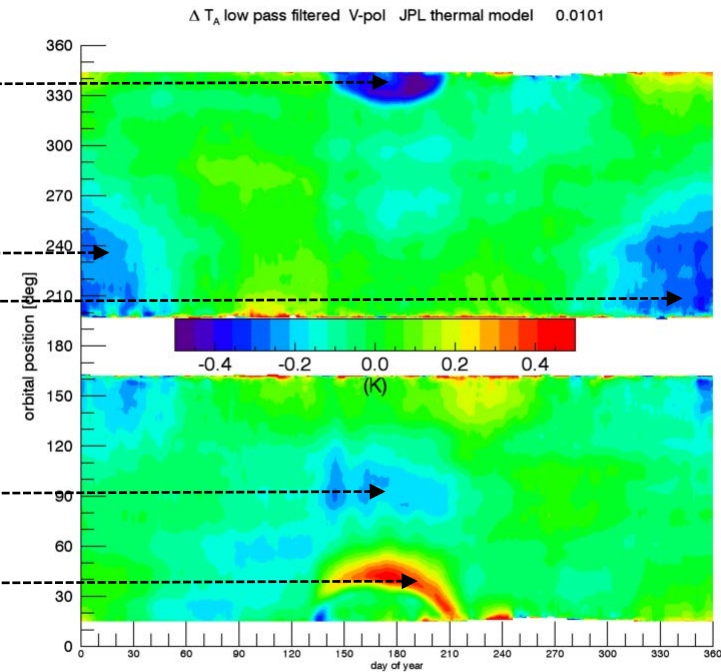
JPL thermal model needs empirical adjustments.

$T_{\text{refl}}$ : JPL thermal model

TA Bias



Eclipse  
Season





# Correction for Land Intrusion

## Sidelobe Correction based on Orbit Simulator

Desired Quantity  
Footprint-Averaged  $T_B$

$$T_{B,3dB} = \frac{\int_{3dB \text{ footprint}} T_{BE,toa}(\bar{\theta}_i) dA}{\int_{3dB \text{ footprint}} dA}$$

Measured Quantity  
 $T_B$  averaged over  $4\pi$  steradians

$$T_A = \frac{1}{4\pi} \int_{4\pi} \mathbf{G}(\mathbf{b}) \Psi(\phi) T_B \frac{\partial \Omega}{\partial A} dA \quad T_{B,mea} = \mathbf{A} \cdot T_A$$

**G**: Antenna gain      **A**: Antenna pattern correction matrix      **Ψ**: Faraday Rotation

Sidelobe Correction

$$\Delta T_B = T_{B,3dB} - T_{B,mea}$$

The 3 dB brightness temperature  $T_{B,3dB}$  and the measured brightness temperatures  $T_{B,mea}$  are found using the on-orbit simulator.

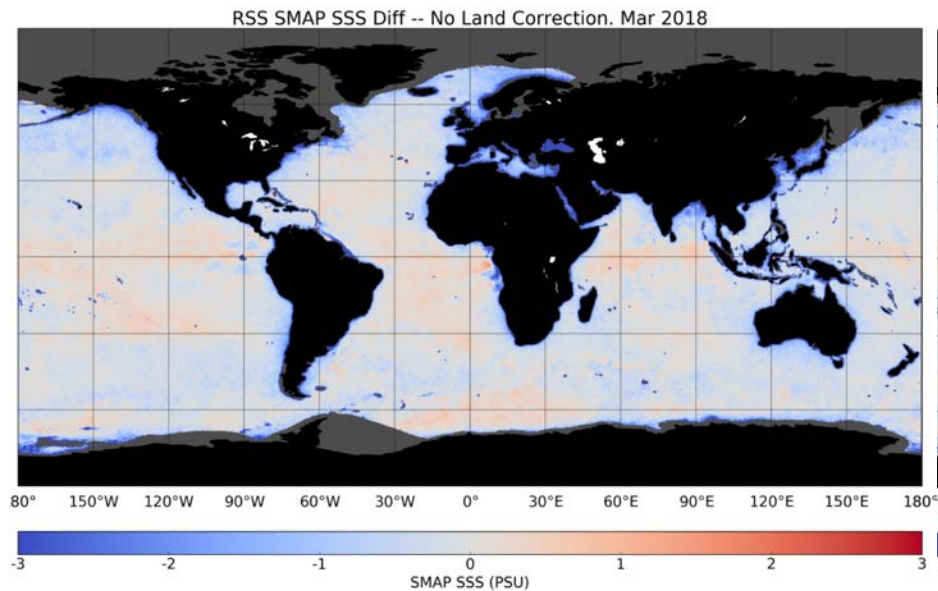
The integration over the Earth's surface is at a spatial resolution of 1 km over the main lobe of the antenna. A coarser resolution is used outside the main lobe.

**Forward Model:** Land emission RTM using climatology of soil moisture + LST (NCEP).

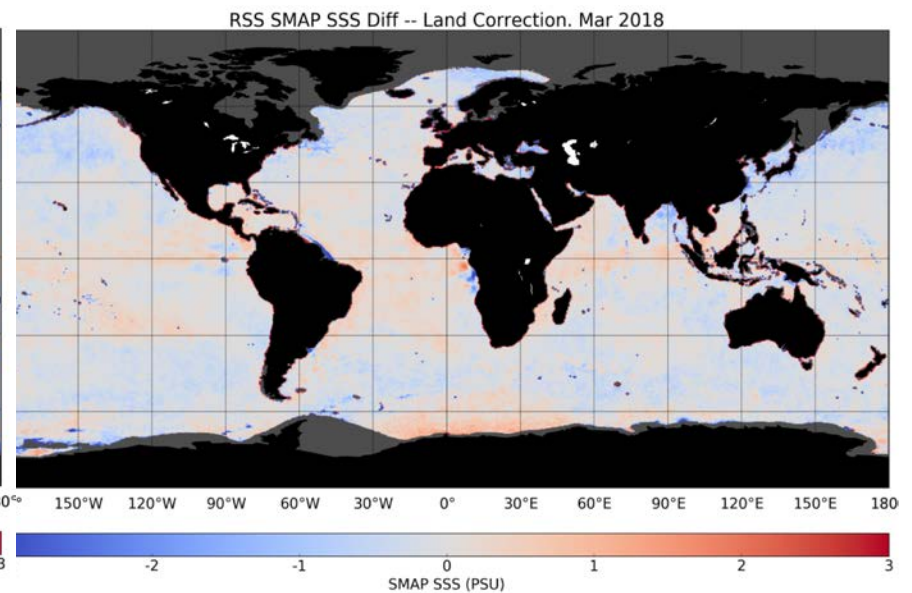
## SMAP V3 Sidelobe Correction Table

polarization	(2 elements, v-pol, h-pol)
cell longitude	(720 elements in 0.5° increment)
cell latitude	(361 elements in 0.5° increment)
ascending/descending	(2 elements)
scan angle	(30 elements in 12° increments)
month	(12 elements)

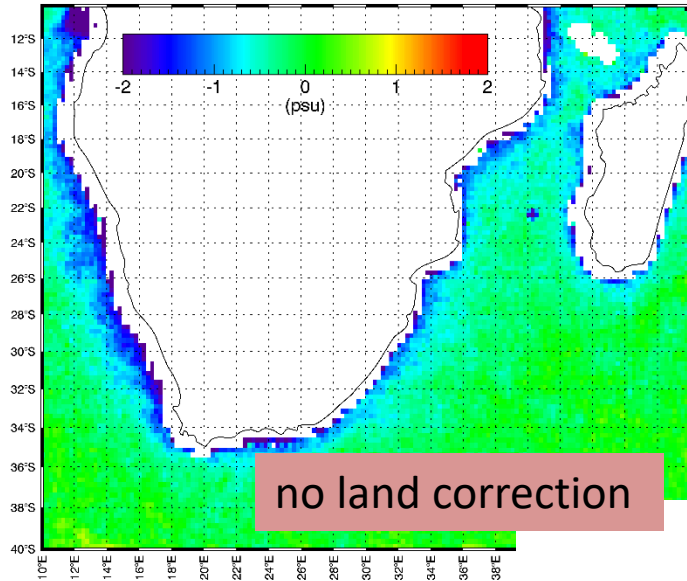
### No Land Correction



### V3.0 Land Correction

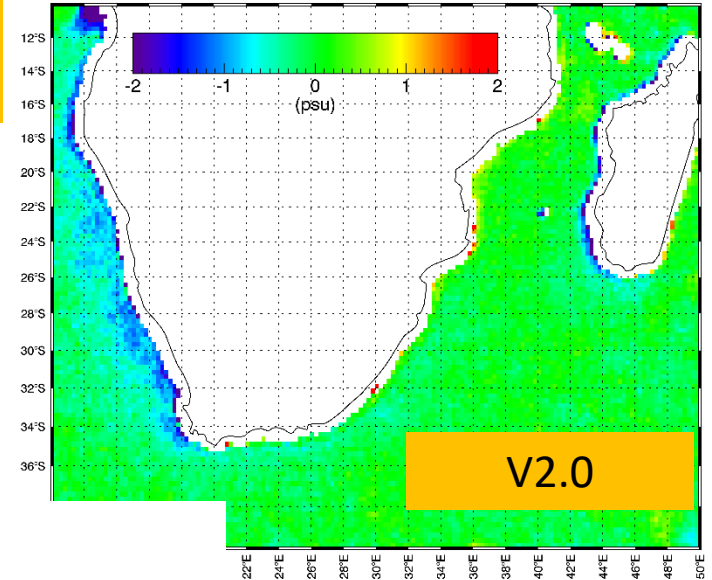


SMAP - HYCOM SSS 2016 01 40km no land correction

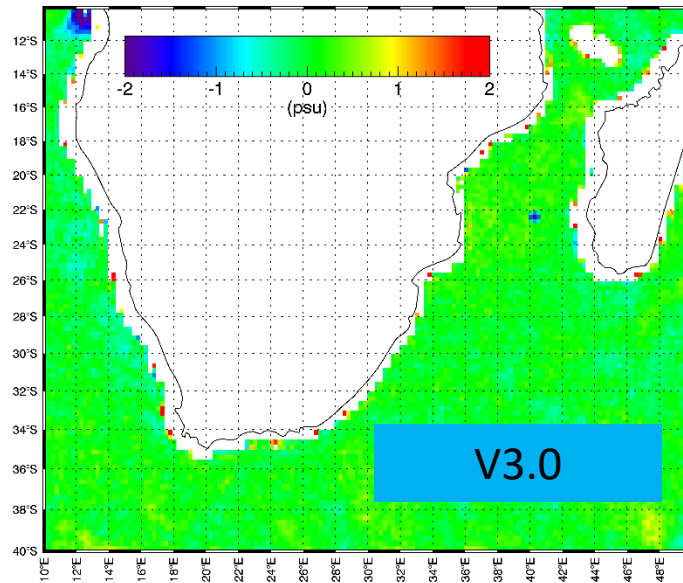


## SMAP – HYCOM L3 monthly

SMAP - HYCOM SSS 2016 01 40km V2.0



SMAP - HYCOM SSS 2016 01 40km V3.0



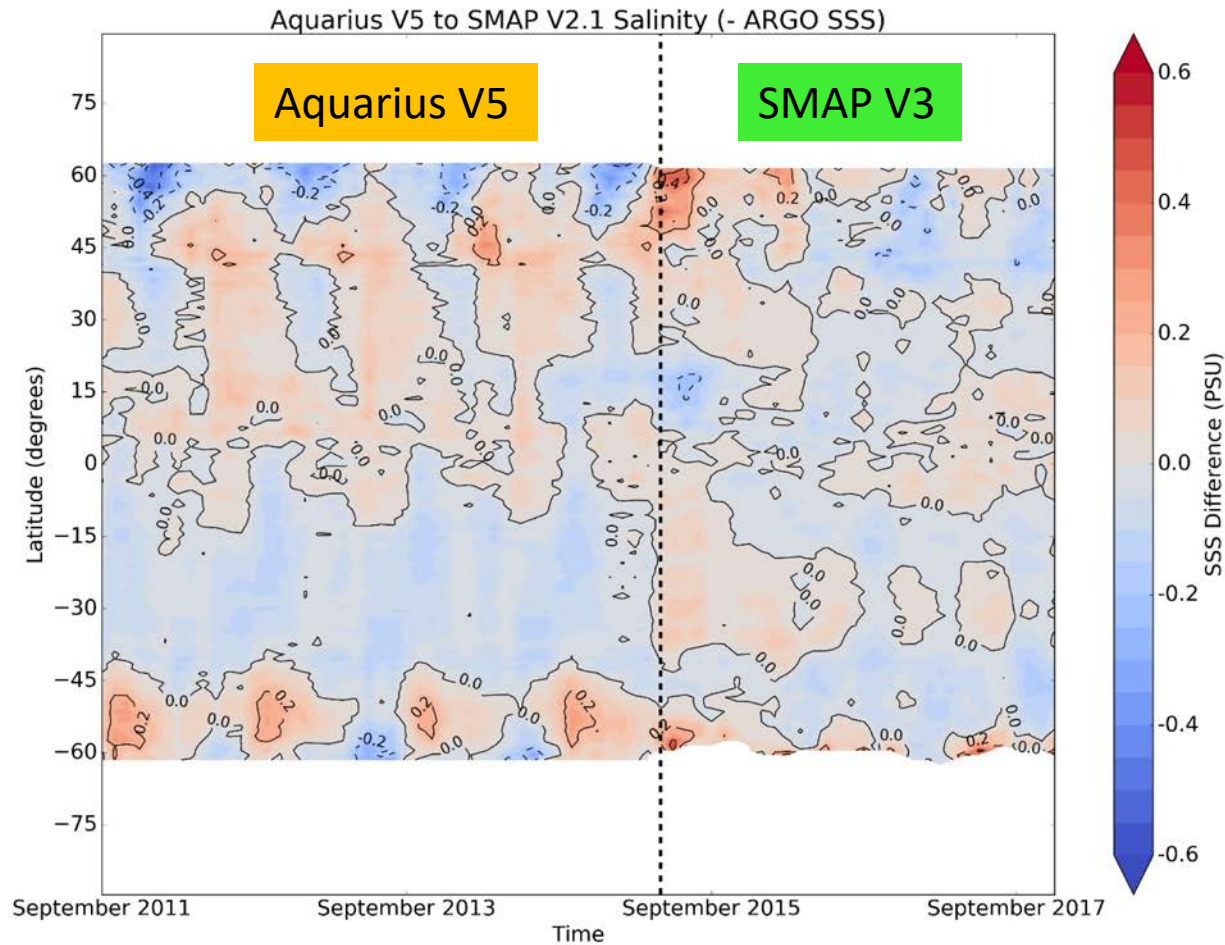
V3.0 land correction  
performs well up to  
about 60 km from the  
coast (1% land  
contamination).

Degrades closer to  
coast.

Work in progress to  
improve.

# Aquarius – SMAP Continuity

use ARGO as common reference

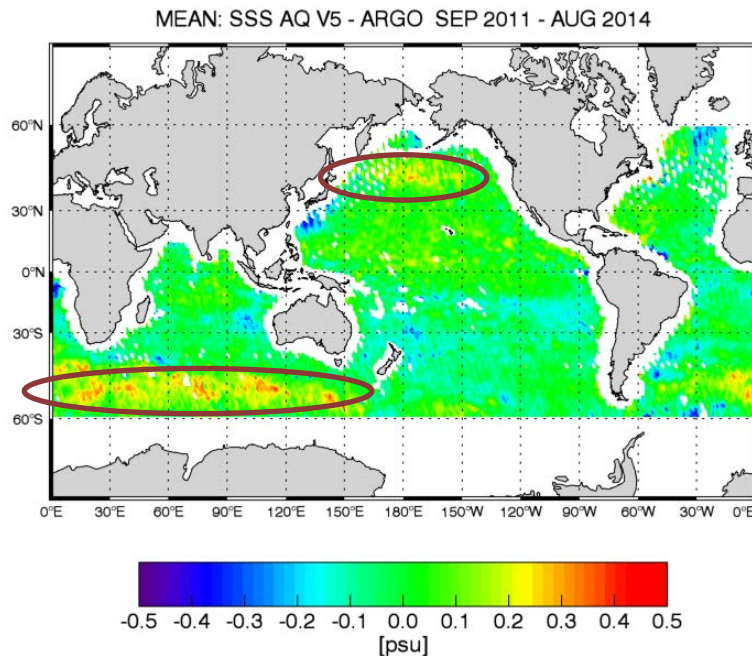




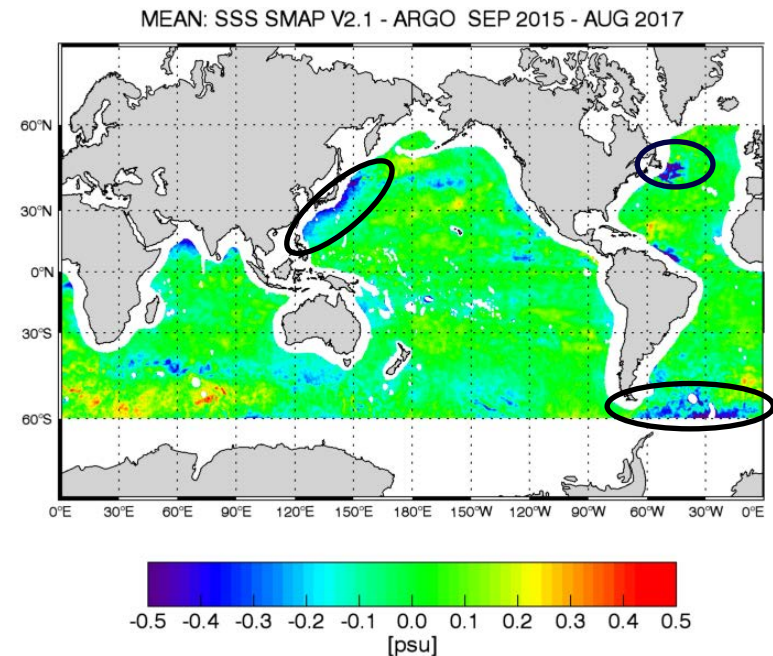
# Aquarius – SMAP Continuity

use ARGO as common reference

Aquarius V5 (2012 – 2014) - ARGO



SMAP V3 (2016 – 2017) - ARGO





# Aquarius – SMAP Continuity

## NASA SCPS

- Residual biases in Aquarius V5 and SMAP V3
- Track down their causes..
  - Wind roughness
  - Galaxy
  - ?
- Differences in sensors and retrieval
  - Ancillary wind speed input
    - Aquarius scatterometer in Aquarius V5 vs CCMP in SMAP V3
  - Different looking geometry
    - Aquarius fixed vs SMAP scanning
    - Directional dependent biases may average out in SMAP.
- Need to keep consistency in GMF for both products



# The Salinity Retrieval Algorithms for the NASA Aquarius Version 5 and SMAP Version 3 Releases

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**Abstract:** The Aquarius end-of-mission (Version 5) salinity data set was released in December 2017. This article gives a comprehensive overview of the main steps of the Level 2 salinity retrieval algorithm. In particular, we will discuss the corrections for wind induced surface roughness, atmospheric oxygen absorption, reflected galactic radiation and side-lobe intrusion from land surfaces. Most of these corrections have undergone major updates from previous versions, which has helped mitigating temporal and zonal biases. Our article also discusses the ocean target calibration for Aquarius Version 5. We show how formal error estimates for the Aquarius retrievals can be obtained by perturbing the input to the algorithm. The performance of the Aquarius Version 5 salinity retrievals is evaluated against salinity measurements from the ARGO network and the HYCOM model. When stratified as function of sea surface temperature or sea surface wind speed, the difference between Aquarius Version 5 and ARGO is within  $\pm 0.1$  psu. The estimated global RMS uncertainty for monthly 100 km averages is 0.128 psu for the Aquarius Version 5 retrievals. Finally, we show how the Aquarius Version 5 salinity retrieval algorithm is adapted to retrieve salinity from the Soil-Moisture Active Passion (SMAP) mission.





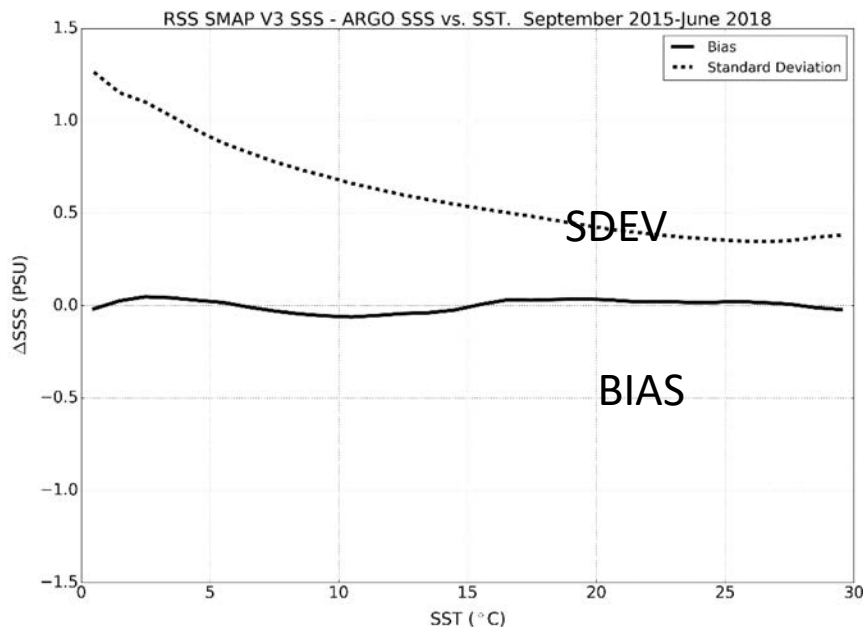
# Backup Slides



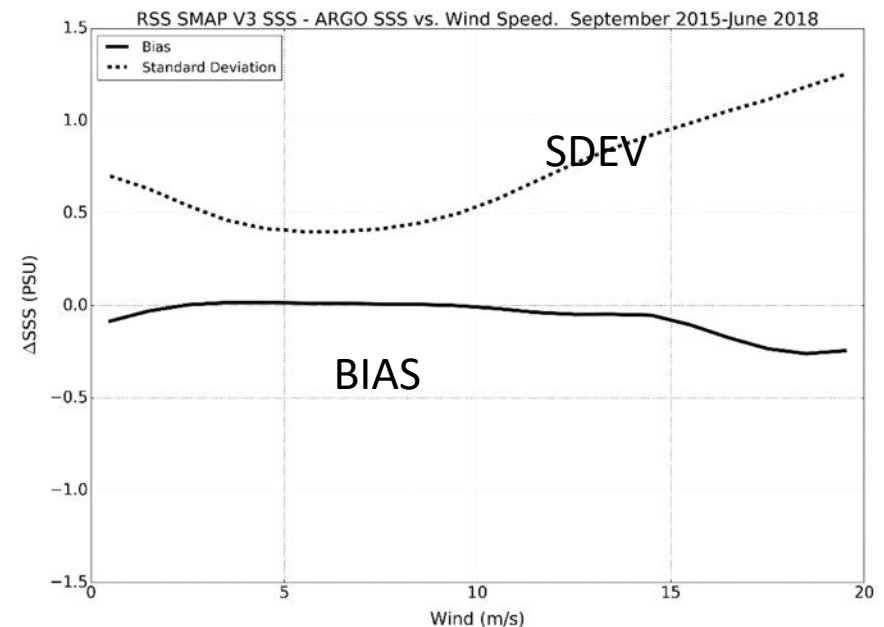
# RSS SMAP V3.0 Level 2 – ARGO

## Open Ocean. Binned v SST and Wind Speed

### VS. SST



### VS. WIND SPEED



No biases over whole SST range and wind speed range (< 15 m/s).  
Consistency of GMF (dielectric model, wind emissivity model).